

Executive Summary

The 2013 Community Summer Study (known as “Snowmass”) brought together nearly 700 physicists to identify the critical research directions for the U.S. particle physics program. This meeting was the culmination of intense work over the last year that defined the most important questions for this field and identified the most promising opportunities to address them. The full written report of the Summer Study is intended to be a key resource for HEPAP, DOE and NSF in setting priorities in particle physics.

To understand the universe, particle physics addresses two main questions: What are the most elementary constituents of nature and what are the forces that cause them to interact?. While these questions are esoteric, finding answers requires significant technical leaps in instrumentation, computing, and accelerators. This development of new technology has continually improved the quality of human life, from industrial techniques, to medical imaging, to computing, and beyond.

The discovery of the Higgs boson in 2012 completes the picture of the particle world called the Standard Model, a remarkable achievement made possible by decades of worldwide collaboration. However, the Standard Model still leaves significant questions unanswered: What is the nature of the Higgs boson? What do neutrino masses tell us? Are the known forces part of a unified structure? What is the composition of 95% of the universe? What mechanism explains the dominance of matter over antimatter? To answer these questions, a variety of approaches is needed rather than a single experiment or technique. Particle physics uses three basic approaches, often characterized as exploration across the cosmic, energy, and intensity frontiers. Each approach employs different tools and techniques, but they ultimately address the same fundamental questions.

These approaches require long-term vision and global partnerships. In designing a program, the priorities of other regions of the world need to be taken into account. The U.S. brings crucial design talent, technology, knowledge and resources that benefit progress in this field regardless of where each experiment is located.

The outline that follows introduces the future directions necessary for further progress in, our understanding of nature. We use the language of “Frontiers” as describing the different experimental approaches as well as for the enabling technologies required to execute the program. Their order does not reflect prioritization.

Particle Physics Frontiers

Intensity Frontier. Experiments at the “Intensity Frontier” explore fundamental questions by using precision measurements to probe quantum effects. They encompass searches for extremely rare processes and for tiny deviations from Standard Model expectations. They have the potential to expose failings in our current understanding, and to discover new laws of physics at very high energies, in many cases exploring beyond the direct reach of high-energy accelerators. This program requires the greatest possible beam intensities, as well as ultra-sensitive, sometimes massive detectors. The Snowmass

47 studies proposed facilities and experiments that will position the US as a global leader in
48 intensity frontier science.

49
50 Neutrinos are the most elusive of the known fundamental particles. Snowmass
51 underscored the reality that we are entering the era of precision neutrino physics in which
52 we can test the three-neutrino paradigm to high accuracy. In light of the recent discovery
53 that the value of θ_{13} is large, there is a clear experimental path forward to determine the
54 properties of neutrinos. The Long-Baseline Neutrino Experiment (LBNE) will measure
55 the mass-hierarchy and is uniquely positioned to determine whether the CP symmetry is
56 violated in the leptonic sector. Future multi-megawatt beams aimed at LBNE, such as
57 those provided from Project X, will be necessary to measure the CP violating phase with
58 sufficient accuracy. An underground LBNE laboratory also allows for the study of
59 atmospheric neutrinos, nucleon decay, and precision measurements of neutrinos from a
60 galactic supernova explosion. This represents a vibrant global program with the US as
61 host.

62
63 Further opportunities to study neutrinos were articulated at Snowmass. An upgrade of
64 the IceCube experiment, provides a promising approach to measure the mass hierarchy
65 using atmospheric neutrinos. Next-generation neutrinoless double-beta decay
66 experiments can reach the sensitivity to determine whether neutrinos are their own
67 antiparticles and are a critical component of a strong neutrino program.

68
69 Flavor observables provide essential probes of new physics. Substantial progress is
70 anticipated this decade with experiments utilizing the Main Injector at Fermilab. These
71 include new measurements of the anomalous magnetic moment of the muon and a deep
72 search for muon-electron conversion. Proposed experiments would probe rare kaon
73 decays to a new level of precision and would retain the US capability to perform heavy
74 quark experiments. Project X will provide the means to probe electric dipole moments at
75 the levels predicted in models of CP violation. Snowmass studies showed that the US
76 can capitalize on current investments and achieve unprecedented sensitivities with the
77 capabilities offered by LHCb at CERN, Belle-II in Japan, and the intense proton source
78 from Project X at Fermilab. Each of these flavor experiments probes a different aspect of
79 new physics in its unique way.

80
81 Dark sectors containing new light weakly coupled particles appear in well-motivated
82 theories. Searches for these particles may proceed with existing facilities, utilizing
83 intense beams and comparatively modest experiments. Studies at Snowmass identified a
84 rich, diverse, and low-cost program that has the potential for high-impact discoveries.

85
86 **Energy Frontier.** The mysteries of the newly discovered Higgs boson were a major theme
87 at Snowmass. The properties of the Higgs boson raise crucial questions that guide large
88 parts of the future particle physics program. These questions call for a three-pronged
89 research program at high energy accelerators: first, to search for new particles with TeV
90 masses predicted by models of electroweak symmetry breaking; second, to make precise
91 measurements of the heavy particles χ_{WZ} , χ_{ZZ} , and the top quark, which can carry the
92 imprint of the Higgs; and, third, to measure the properties of the Higgs boson itself to

93 very high precision. Questions about the Higgs boson also inspire the search for the dark
94 matter particles and for flavor-changing rare decays, since in both cases, the motivating
95 theory often comes from models of the Higgs and its role in symmetry-breaking.

96
97 For at least the next fifteen years, the experiments at the Large Hadron Collider at CERN
98 will drive the energy frontier program forward. Especially in its high-luminosity phase
99 the LHC is expected to explore deeply for new particles produced through either the
100 strong or the electroweak interactions. The LHC will study rare decays using a sample of
101 billions of top quarks and probe for new dynamics of W , Z , and Higgs at TeV
102 energies. It will measure Higgs boson couplings at the few-percent level and provide the
103 first measurement of the Higgs self-coupling. The LHC experiments have already proven
104 their ability to work as global collaborations. Technology, insights, and leadership from
105 the US have played indispensable roles in these experiments.

106
107 There is strong scientific motivation for continuing this program with lepton colliders.
108 Experiments at lepton colliders allow searches for new particles with unequivocal
109 discovery or exclusion, complementing those at the LHC. They can improve the precision
110 of our knowledge of the W , Z , and top properties by an order of magnitude,
111 potentially bringing these measurements into confrontation with theory. They can reach
112 sub-percent precision in the Higgs boson properties, allowing discoveries of percent-level
113 deviations predicted in theoretical models. A global effort has now completed the
114 technical design of the International Linear Collider (ILC), an accelerator that will
115 provide these capabilities. The Japanese high energy physics community has named this
116 facility as its first priority.

117
118 The Snowmass study considered many other options for high energy colliders that might
119 be realized over a longer term. These included linear and circular e^+e^- colliders, muon
120 colliders, and photon colliders. Serious study was begun on the physics program of
121 future proton colliders at 33 and 100 TeV center of mass energy.

122
123 **Cosmic Frontier.** Experiments at the “Cosmic Frontier” include innovative,
124 interdisciplinary approaches to solving the mysteries of dark matter and dark energy,
125 which are the dominant components of the universe but whose fundamental nature is
126 almost completely unknown.

127 The Snowmass process produced a clear articulation of how the different approaches to
128 dark matter – direct detection, indirect detection, accelerator-based searches, simulations,
129 and astrophysical surveys – each provide unique and necessary information, as well as a
130 census of present-day and proposed facilities and their capabilities. The leaps in
131 sensitivity of the new facilities bring us to a realm in which major discoveries could be
132 imminent.

133 Snowmass also strongly reinforced the roles that cosmic surveys play in particle physics.
134 Stage III and Stage IV dark energy imaging and spectroscopic surveys will shrink the
135 errors as recommended in previous community studies, but they will do even more for
136 particle physics: the richness of the data and detailed attention to systematic error
137 management will enable many new tests of the behavior of dark energy and General

138 Relativity over a wide range of distance scales and settings. Cosmic microwave
139 background (CMB) experiments will probe the physics of Inflation, with sufficient
140 sensitivity to falsify significant classes of models.

141 Remarkably, future cosmic surveys, as well as future polar-ice neutrino projects, will also
142 provide precise information about neutrino properties, including the mass hierarchy, the
143 number of light neutrinos, and the sum of the masses of the neutrino species. Combining
144 this information with accelerator-based and reactor-based neutrino experiments, as well
145 as other experiments such as those searching for neutrinoless double-beta decay, will
146 accelerate our understanding of fundamental neutrino properties and enable us to derive
147 meaning from potential inconsistencies.

148 Finally, the Snowmass process reiterated the unique information we can gain from
149 studies of cosmic particles, including detection of GZK neutrinos, an extremely high-
150 energy flux of neutrinos that will enable the study of interactions at energies of 100 TeV,
151 and detection of significant numbers of the highest energy cosmic rays produced in
152 nature. Essential technologies and facilities, and required advances in theory, were
153 identified for all these areas

154 The largest projects are, appropriately and necessarily, international. The U.S. is still the
155 leader in this quickly evolving area, but other regions with intensive interest in this
156 physics are advancing rapidly.

157 **Theoretical Physics:** Progress in particle physics requires interplay between theory and
158 experiment. The US has been the world leader for many decades in particle theory, and a
159 sustained strong and vibrant program remains essential for the success of US particle
160 physics. Theory has been a driving force both in the development and testing of the
161 Standard Model, including the Higgs discovery. Theory has been critical in formulating
162 the big questions in the field, setting out hypotheses that address them, and proposing
163 experimental strategies to confirm or refute them. At the same time, theorists seek new
164 structures that might provide unanticipated resolutions. The success of the US program
165 has rested on principal investigators in universities and national labs, working with
166 postdoctoral fellows and graduate students, and collaborating with both US and
167 international researchers.

168

169 *Enabling Frontiers*

170 **Capabilities Frontier.** Accelerators for the widely diverse range of next generation
171 intensity frontier experiments have the common requirements of delivering multi-MW
172 proton beams with a flexible, experiment-dependent timing structure. Such requirements
173 are beyond the capabilities of any present accelerator. The proposed Project X, when
174 fully realized, would deliver such beams over a broad energy range from 0.25 to 120
175 GeV. It would also be a platform for future muon facilities with muon storage rings as a
176 possible first step toward neutrino factories. Innovative cyclotrons for neutrino studies
177 may be part of this program. Critical technical issues that must be addressed for all

178 proposed intensity frontier accelerators include controlling beam loss and understanding
179 the limits of material damage in high power targets.

180
181 The currently proposed energy frontier accelerators incorporate significant US
182 contributions to the accelerator technology, both for the LHC and for the ILC. US
183 physicists are involved in the upgrades of the LHC accelerator through the LHC
184 Accelerator Research Program (LARP). The technology of Nb₃Sn superconducting
185 magnets, pioneered in the US and developed for application through LARP, has a key
186 role in the high-luminosity proposal. The ILC design embodies US leadership
187 contributions in beam dynamics, superconducting RF technology, damping ring design,
188 and beam delivery. This design is ready to proceed to construction.

189
190 Innovative concepts for higher energy accelerators, both proton-proton and lepton
191 colliders, to be realized on much longer time scales, were also discussed. The CLIC two-
192 beam accelerator is based on X-band, warm linac technology, with a 50 km length to
193 reach 3 TeV. A second approach is to cool and collide muon beams. This muon collider
194 – if proven feasible – would fit a 3 TeV collider onto the present Fermilab site. The
195 muon beam approach has a very large overlap with capabilities needed for intensity
196 frontier accelerators. A third approach would use wakefields driven either by beams or
197 lasers to achieve accelerating fields of 10 to 100 GeV per meter. A hadron collider of
198 higher energy can be designed within the development reach of existing materials and
199 technologies for a tunnel of 100 km or larger. Higher energy or lower cost would require
200 development of new classes of superconducting magnets. US scientists are currently
201 leading the development of high field magnets, the key enabling technology in the quest
202 for higher energies.

203
204 Maximizing the potential of any next energy frontier accelerator will require continuing
205 an integrated multi-laboratory engineering program carrying out innovative long-range
206 research. The US national laboratories, with their unique expertise, should invest in the
207 development of technologies for accelerators capable of higher energies, more luminous
208 beams, and more efficient operation that will provide the setting for future high energy
209 physics experiments. This program of R&D and accelerator stewardship is likely to bring
210 benefits to all areas of science and industry that make use of accelerator technology.

211
212 Many experiments searching for dark matter, proton decay or seeking to determine the
213 properties of neutrinos must be located underground to shield the sensitive experiments
214 from cosmic ray backgrounds. Underground facilities are located or proposed in North
215 and South America, Europe, Asia and in the Antarctica ice. About 1,000 particle
216 physicists from the United States (and more from outside the United States) are currently
217 taking part in experiments in underground facilities around the world. The scope of
218 underground capabilities in all regions is expected to increase by the end of the decade to
219 accommodate increased experimental demand. Locating the Long Baseline Neutrino
220 Experiment (LBNE) underground in South Dakota would allow this experiment to realize
221 its full scientific potential and could make it an anchor for possible future domestic
222 underground capabilities for a broader range of compelling experiments.

223 **Instrumentation Frontier.** The “Instrumentation Frontier” enables experiments to answer

224 the science questions described above. High Energy Physics has a long and distinguished
225 history of inventing, designing and building the specialized instrumentation required for
226 its experimental research. At Snowmass, the vision for a US instrumentation program for
227 particle physics was formulated, which would enable the US to maintain a scientific
228 leadership position in a broad, global, experimental program. The program's strategy is to
229 identify from the plethora of technologies those that will overcome key barriers to
230 answering the science questions. Some key technology areas for further investment have
231 been identified at Snowmass.

232 Executing this strategy requires integrating the diverse capabilities and resources of
233 universities, national laboratories, other branches of science, and industry into detector
234 R&D collaborations, which would reestablish the importance of innovation through a
235 domestic instrumentation development program. A coordinating panel will help articulate
236 the mission of this program and facilitate its implementation. The goals of the program
237 are to develop both incremental and transformational, cost-effective technologies with
238 maximal scientific reach, based on the technological strengths in the US. A stably and
239 adequately funded generic instrumentation program will ensure that particle physics
240 invests in its future and establishes a foundation for a competitive, healthy program in the
241 long term.

242 **Computing Frontier.** Computing is a major component of all particle physics experiments
243 and in many areas of theoretical physics. The “Computing Frontier” was divided into a
244 number of subgroups covering user needs and infrastructure. The user needs groups
245 covered each of the experimental frontiers and four theoretical areas. The infrastructure
246 groups examined trends in computing to predict how technology will evolve and how it
247 will affect the costs and capabilities of computing systems, networks and storage.

248 A number of concerns were identified. Given the changes in chip technology and
249 evolution of high performance systems to include multi-core chips and accelerators, how
250 will we write the parallel codes that will be needed in the future and how will we train the
251 personnel to develop, support and maintain them? Upgrades to the LHC will place more
252 demands on the distributed computing systems for ATLAS and CMS. Attention needs to
253 be paid to wide-area networking trends to ensure that the network does not become a
254 bottleneck in the future. It would be desirable to better coordinate software development
255 for a number of intensity frontier experiments whose needs are not at the level of LHC
256 experiments, but will nevertheless become substantial. The price of storage and
257 bandwidth of disks may be of significant concern in the future. Prices may not drop as
258 rapidly as in the past and, while disks have become larger, their interfaces have not
259 become significantly faster. Early attention to these issues has the potential to increase
260 efficiency, reduce costs, enable significantly more realistic theoretical calculations and
261 avoid computing bottlenecks in the experimental program that could limit scientific
262 progress.

263 *Communication, Education and Outreach*

264
265 The particle physics community recognizes the critical importance of consistent and
266 coherent communication, education and outreach. These foster nationwide support for

267 the field, develop the next generation of physicists and ensure scientifically literate
268 citizens. More physicists must engage in CE&O activities to translate the American
269 public's fascination for particle physics research into the support necessary to enable the
270 field to answer its biggest questions. Existing activities must be augmented with
271 dedicated personnel who will enhance these efforts, provide nationwide coordination and
272 spearhead new initiatives. Such actions include a central communication, education and
273 outreach office for physicists, materials designed to inform the public about direct and
274 indirect applications of particle physics research, sustainable methods to collect statistics
275 on workforce development and technology transfer, professional development
276 opportunities for educators and new learning opportunities for students of all ages.

277

278

279 High-energy physicists are pioneers of large scientific projects. The Tevatron and now
280 the Large Hadron Collider give examples of projects that were developed over 25 years,
281 with global collaboration, that led to crowning discoveries. As these projects were
282 developed, we have evolved a scientific community that links together scientists from all
283 over the world pursuing common goals. With the program outlined above, we will
284 achieve more successes in the future in our common quest to gain a fuller understanding
285 of the universe.

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292 Big Questions for Particle Physics

293

294 Through the previous six decades of precision and discovery-level particle physics, we
295 have learned much about the basic laws that govern the universe. We have uncovered
296 the laws that describe the subnuclear forces and, with the discovery of the Higgs boson,
297 the agent that we believe should give mass to all elementary particles. However, there is
298 still much that we do not understand. The advance in our knowledge of elementary
299 particle physics has sharpened the questions in that domain. Recent discoveries about the
300 matter and energy content of the universe have brought new questions that are equally
301 fundamental, and equally vexing.

302

303 One of the goals of Snowmass was to develop a framework of scientific questions that
304 can form the basis for a future program in particle physics. To introduce a summary of
305 the results of Snowmass, we propose a set of ten Particle Physics Questions. The search
306 for their answers will be carried out with a broad range of experimental methods, cutting
307 across the "frontiers" around which the Snowmass study was organized.

308

309 1. How do we understand the Higgs boson? What principle determines its couplings to
310 quarks and leptons? Why does it condense and acquire a vacuum value throughout the

311 universe? Is there one Higgs particle, or many? Is the Higgs particle elementary or
312 composite?
313

314 2. How do we understand the multiplicity of quarks and leptons? What principle
315 determines their masses and mixings? Why is the mixing pattern apparently different for
316 quarks and leptons? Why is the CKM CP phase nonzero? Is there CP violation in the
317 lepton sector?
318

319 3. How do we understand the neutrinos? Why are their masses so different from those of
320 other fermions? Are neutrinos their own antiparticles? Are their small masses connected
321 to the presence of a very high mass scale? Are there new interactions invisible except
322 through their role in neutrino physics?
323

324 4. How do we understand the matter-antimatter asymmetry of the universe? Why are the
325 interactions of particles and antiparticles not exactly mirror opposites? What principle
326 determines the excess of baryons over antibaryons that we see in the universe?
327

328 5. How do we understand the substance of dark matter? This is the dominant component
329 of mass in the universe. What is the dark matter made of? Is it composed of one type of
330 new particle or several? What principle determined the current density of dark matter in
331 the universe? Are the dark matter particles connected to the particles of the Standard
332 Model?
333

334 6. How do we understand the dark energy? Is it a static energy per unit volume of the
335 vacuum, or is it dynamical and evolving with the universe? What principle determines
336 its value?
337

338 7. How do we understand the origin of structure in the universe? The inflationary
339 universe model requires new fields active in the early universe. Where did these come
340 from, and how can we probe them today?
341

342 8. How do we understand the multiplicity of forces? Are there additional forces that we
343 have not yet observed? Are there additional quantum numbers associated with new
344 fundamental symmetries? Are the four known forces unified at very short distances?
345 What principles are involved in this unification?
346

347 9. Are there new particles at the TeV energy scale? Such particles are motivated by the
348 problem of the Higgs boson, and by ideas about spacetime symmetry such as
349 supersymmetry and extra dimensions. If they exist, how do they acquire mass, and what
350 is their mass spectrum? Do they carry new sources of quark and lepton mixing and CP
351 violation?
352

353 10. Are there new particles at higher energies? Are there new invisible particles? Such
354 particles are motivated by the strong CP problem, the problem of neutrino mass, the
355 problem of the origin of cosmic inflation, the question of grand unification, and the

356 question of unification of microscopic forces with gravity. Can we learn the nature of
357 these particles from our experiments?

358

359 The search for answers to these questions is intimately tied to the development of
360 technology. High energy physics experiments and high energy accelerators put
361 extraordinary demands on sensors, precision engineering, and data management,
362 incorporated into devices of very large scale. Our community invents new technologies
363 to address these needs and brings them into robust use. The progress of our field
364 requires both technology development directed at the problems of specific experiments
365 and the development of new technologies that allow leaps to higher performance or
366 decreased cost for devices of broad application. This technology development for
367 accelerators and detectors ultimately benefits all of physical science.

368

369 In many areas of physics experimentation, there are specific technological developments
370 that would be of enormous benefit. Existing technologies are unlikely to meet the science
371 needs of future high energy physics experiments. New technologies need to be explored
372 that could lead to transformative advances, enabling cost-effective high energy physics
373 experiments but also new initiatives of broad importance.

374

375 1. As experiments continue to reach for rarer processes, more precise measurements,
376 higher energies and luminosities, and more inclusive observations, they place increasing
377 demands on detector technology. How do we achieve the detectors of finer granularity,
378 larger volume, greater radiation hardness, lower cost, and higher speed that will in large
379 part determine our experimental reach?

380

381 2. Paradigm altering technology developments are occurring in electronics and materials
382 design, potentially offering breakthrough capabilities. How can these advances be
383 incorporated into new detectors with improved overall performance? How do we make
384 best use of the resources available in universities, national laboratories and industry to
385 develop new detector systems?

386

387 3. What technologies will be needed to acquire, analyze and store the enormous amounts
388 of data from future experiments? Can local intelligence be incorporated to manage data
389 flow? How will we fully and efficiently utilize data stored in large databases?

390

391 4. Scaling of current accelerator designs to higher energy leads to machines of very large
392 size, cost, and power demand. Can new technologies lead to more practical
393 strategies? Is there an ultimate highest energy for colliders?

394

395 5. Proposed experiments both at low and at high energy call for particle beams of
396 extreme brightness. Are there technologies to achieve high beam power in a better
397 controlled and more cost-effective way?

398