

1 **Planning the Future of U.S. Particle**
2 **Physics**

3 **Report of the 2013 Community Summer Study**

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Foreword

11

12 Particle physics research in the United States is a vibrant field, with discoveries stretching back more than
 13 half a century. Over the past year, more than 1000 U.S. particle physicists engaged in an in-depth process to
 14 define the most important questions for our field and identify the most promising opportunities to address
 15 these questions.

16 The process had its roots in a series of meetings held at Snowmass, Colorado over the past thirty years to
 17 take stock of progress in particle physics and chart the field’s future. The last such periodic meeting was
 18 held in 2001. In 2011 the Division of Particles and Fields of the American Physical Society decided that
 19 the time was ripe for a new community study. Preparatory meetings during 2012 and 2013 began with a
 20 Community Planning Meeting at Fermilab, October 11-13, 2012. A final plenary meeting was held at the
 21 University of Minnesota, July 29 - August 6, 2013.

22 The 2013 Community Summer Study (the “Snowmass study”) was designed to enumerate the questions the
 23 particle physics community wishes to answer over the next two decades, and plan how to answer them. The
 24 study does not prioritize activities, but aims to ask and answer hard questions. A subsequent prioritization
 25 panel with broad community representation will place these questions and answers within realistic budgetary
 26 scenarios. The study has produced this resource book, charting aspirations of the U.S. particle physics
 27 community over the next ten to twenty years, for use by the panel in its deliberations. We also intend the
 28 report to convey the health and diversity of the U.S. program, in a global context, to our colleagues and
 29 fellow citizens.

30 Although we found it convenient to retain the “frontier” categories of the previous Particle Physics Project
 31 Prioritization Panel (“P5”), whose last report was issued in 2008, the division of the field into such categories
 32 should not obscure the focus on fundamental questions of physics, which, by their nature, cross such frontiers.
 33 These inter-frontier discussions have been a major component of the Minnesota meeting.

34 This volume can be read at various levels. An executive summary precedes a more detailed summary
 35 chapter. Each frontier (Intensity, Energy, Cosmic, Theory, Capabilities, Instrumentation, Computing, and
 36 Communication) has its own chapter containing further details. Reference is made to submissions by each
 37 frontier’s subgroups, and to contributed white papers.

38 We thank all the host institutions and organizers of the preparatory and Minnesota meetings for their efforts
 39 on behalf of this review of our field.

40 For the Conveners and the DPF Executive Committee,

	Jonathan L. Rosner	DPF Chair
	Ian Shipsey	DPF Chair-Elect
41	Nick Hadley	DPF Vice-Chair
	Pierre Ramond	DPF Past Chair

Executive Summary

42

43 The 2013 Community Summer Study, known as “Snowmass,” brought together nearly 700 physicists to
44 identify the critical research directions for the United States particle physics program. Commissioned by the
45 American Physical Society, this meeting was the culmination of intense work over the past year by more than
46 1000 physicists that defined the most important questions for this field and identified the most promising
47 opportunities to address them. This Snowmass study report is a key resource for setting priorities in particle
48 physics.

49 Particle physicists seek to understand the structure of the universe. We address two main questions: What
50 are the most elementary constituents of nature, and what are the forces that cause them to interact?
51 These questions are fundamental, and the desire to explore them is a defining characteristic of the human
52 spirit. At the same time, finding the answers has practical value: It helps drive technical innovation in
53 instrumentation, computing, and accelerators, and leads to the development of a skilled technical workforce.
54 The development of new technologies, from industrial techniques, to medical imaging, high-performance
55 computing, and beyond, has continually improved the quality of human life.

56 The discovery of the Higgs boson in 2012 was a remarkable achievement made possible by decades of
57 worldwide collaboration. The existence of the Higgs boson was predicted in the 1960s. By last year it
58 was the sole missing piece of the theory we call the Standard Model. This theory provides a coherent picture
59 of the strong, weak, and electromagnetic interactions, with the latter two unified in an “electroweak” theory.
60 The Standard Model contains quarks, leptons, force carriers, and now the Higgs boson.

61 However, the Standard Model still leaves significant questions unanswered. What is the nature of the Higgs
62 boson? What can we learn from discovering that neutrinos have mass? Can the known forces be further
63 unified? The particles of the Standard Model make up only 5% of the universe — what is the other 95%?
64 Why is the universe almost all matter and no antimatter?

65 Many different techniques are needed to answer these questions. Particle physics explores three basic
66 frontiers, often called cosmic, energy, and intensity. We continue that designation in this report. Each
67 frontier uses its own tools and techniques, but all ultimately address the same fundamental questions.

68 The experiments that address these questions are ambitious, large-scale projects. Mounting them requires
69 long-term vision. We are fortunate that our priorities are shared by physicists in other regions of the world,
70 so that these experiments can be realized as global partnerships. The U.S. brings crucial leadership, design
71 talent, technology, and resources that will be essential to these experiments wherever they are located.

72 The outline that follows introduces the future directions necessary for further progress in our understanding
73 of nature at its most fundamental level. The order does not reflect prioritization.

74 *Particle Physics Frontiers*

75 **Intensity Frontier:** Experiments at the “Intensity Frontier” explore fundamental questions by using
76 precision measurements to search for extremely rare processes and for tiny deviations from expectations.
77 They can reveal new laws of physics at very high energies, in many cases exploring beyond the direct reach
78 of high-energy accelerators. They often require the greatest possible beam intensities and ultra-sensitive
79 detectors. The Snowmass study identified facilities and experiments that will ensure the global leadership
80 of the U.S. in Intensity Frontier science.

81 Neutrinos are the most elusive of the known fundamental particles. We know of three species, which can
82 change (“oscillate”) into one another under the right conditions. We have recently discovered that one of the
83 neutrino oscillation parameters is large enough to let us measure two fundamental properties of neutrinos.

84 First, we can hope to determine which of the three neutrinos is heaviest and which is lightest (the mass
85 hierarchy). Second, we can determine if leptons violate CP (charge-parity) symmetry, one of the most
86 important symmetries of nature. We have known for almost fifty years that CP is violated by quarks; now
87 the focus has shifted to leptons.

88 The Long-Baseline Neutrino Experiment (LBNE) will measure the mass hierarchy and is uniquely positioned
89 to determine whether leptons violate CP. Future multi-megawatt beams aimed at LBNE, such as those from
90 Project X at Fermilab, would enable studies of CP violation in neutrino oscillations with conclusive accuracy.
91 An underground LBNE detector would also permit the study of atmospheric neutrinos, proton decay, and
92 precision measurement of any galactic supernova explosion. This represents a vibrant global program with
93 the U.S. as host.

94 The Snowmass study discussed further opportunities to learn about neutrinos. An upgrade of the IceCube
95 experiment, involving photodetectors buried in Antarctic ice, could provide a promising approach to measure
96 the mass hierarchy using atmospheric neutrinos. Next-generation neutrinoless double-beta decay experiments
97 could reach the sensitivity necessary to determine whether neutrinos are their own antiparticles, and are a
98 critical component of a strong neutrino program.

99 Transitions among different types, or “flavors,” of quarks and leptons provide essential probes of new
100 physics, and are a central element of the Intensity Frontier. Substantial progress toward understanding
101 these transitions can be made in this decade with experiments utilizing the Fermilab accelerator complex.
102 These include a new measurement of the anomalous magnetic moment of the muon, a sensitive search for
103 muon-to-electron conversion, and a proposed experiment to probe rare K meson decays to a new level of
104 precision.

105 If particles have a property known as electric dipole moments (EDMs), they will violate CP symmetry.
106 Finding EDMs would have profound implications, allowing us to check models of CP violation and to help
107 understand the preponderance of matter in the universe. Project X could provide the means to make incisive
108 EDM measurements of unprecedented precision.

109 The U.S. has also made substantial investments in other flavor physics experiments. Snowmass studies
110 showed that the U.S. can capitalize on them, achieving unprecedented sensitivities with the capabilities
111 offered by the LHCb experiment at CERN, Belle-II and J-PARC in Japan, and BESIII in China. Each
112 probes a different aspect of new physics in a unique way. Project X will complement and extend results of
113 these experiments.

114 New light, weakly coupled particles appear in many theoretical models, especially those of the cosmic dark
115 matter. Some searches for these particles are feasible with intense beams and comparatively modest detectors
116 at existing facilities. The Snowmass studies identified a rich, diverse, and low-cost program with a potential
117 for high-impact discoveries, illustrating the importance of modest-scale experiments to complement large-
118 scale efforts.

119 **Energy Frontier:** Experiments at the “Energy Frontier” make use of high-energy colliders to directly
120 produce heavy elementary particles and to search for new ones. The properties of the Higgs boson raise
121 crucial questions that guide large parts of the future particle physics program. The Higgs boson discovery
122 calls for a three-pronged research program at high-energy accelerators, to: (1) determine the properties of
123 the Higgs boson as accurately as possible, (2) make precise measurements of the heavy particles W , Z , and
124 the top quark, which can carry the imprint of the Higgs field, and (3) search for new particles predicted
125 by models of the Higgs boson and electroweak symmetry breaking. These topics also overlap with those in
126 other frontiers. The expectation of particles with masses at or above one TeV directly motivates searches
127 for dark matter particles and flavor-changing rare decays.

128 For at least the next fifteen years, the experiments at the Large Hadron Collider (LHC) at CERN will
129 drive the Energy Frontier program forward. The LHC experiments are successful global collaborations. U.S.
130 leadership and contributions to detector and accelerator components, technology, and physics insight have
131 been indispensable to that success.

132 Following the Higgs boson discovery, the LHC moves into a phase of precision study of the properties of
133 this particle. The high-luminosity LHC will measure Higgs boson interactions at the few-percent level. LHC
134 operation at higher energy over the next ten years, and, later, at higher luminosity, will continue the search
135 for new particles produced through the strong or electroweak interactions. It will probe for new dynamics
136 of the W , Z , and Higgs bosons at TeV energies and study rare decays using billions of top quarks.

137 Compelling science motivates continuing this program with experiments at lepton colliders. Experiments at
138 such colliders can reach sub-percent precision in Higgs boson properties in a unique, model-independent way,
139 enabling discovery of percent-level deviations from the Standard Model predicted in many theories. They
140 can improve the precision of our knowledge of the W , Z , and top quark well enough to allow the discovery
141 of predicted new-physics effects. They search for new particles in a manner complementing new particle
142 searches at the LHC. A global effort has completed the technical design of the International Linear Collider
143 (ILC) accelerator and detectors that will provide these capabilities in the latter part of the next decade. The
144 Japanese particle physics community has declared this facility as its first priority for new initiatives.

145 The Snowmass study considered many other options for high-energy colliders that might be realized over
146 a longer term. These included higher-energy linear colliders, circular e^+e^- colliders, muon colliders, and
147 photon colliders, all of which merit continued study. The Snowmass study identified, in particular, the
148 promise of a 100 TeV-class hadron collider (VLHC), which would provide a large step in energy with great
149 potential for new insights into electroweak symmetry breaking and dark matter. The feasibility of such a
150 machine should be clarified through renewed accelerator R&D and physics studies over the next decade.

151 In all of the above projects, U.S. leadership in developing detector and accelerator technologies is playing a
152 critical role. These U.S. initiatives are essential to meet the world-wide scientific goals in particle physics.

153 **Cosmic Frontier:** We now know that 95% of the universe is in the form of dark matter and dark energy.
154 These components are responsible for the structures and accelerated expansion of the universe observed
155 today, but their fundamental nature is almost completely unknown. “Cosmic Frontier” experiments are
156 designed to determine the nature of dark matter and dark energy and to use the universe as a laboratory to
157 search for new fundamental particles and interactions. Along with the other frontiers, the Cosmic Frontier
158 provides particle physics with clear evidence for physics beyond the Standard Model, profound questions of
159 popular interest, frequent new results and surprises with broad impacts, many opportunities for discovery
160 with unique probes, important cross-frontier topics, and a full range of project scales providing flexible
161 programmatic options.

162 Our studies showed how different approaches to dark matter — direct detection; indirect detection of gamma
163 rays, neutrinos, and antimatter; accelerator-based searches; simulations; and astrophysical surveys — provide
164 unique and necessary information. A census of current and proposed experiments and their capabilities shows
165 that new direct and indirect detection experiments will probe dark matter masses inaccessible to colliders
166 and provide leaps in sensitivity at moderate cost. For many leading dark matter candidates, including
167 well-known examples of weakly-interacting massive particles and axions, this progress will lead to the first
168 non-gravitational signals of the dark universe and open the door to the identification of dark matter, with
169 far-reaching implications for both particle physics and astronomy.

170 Snowmass studies also strongly reinforced the roles that cosmic surveys play in particle physics. Current and
171 upcoming dark energy imaging and spectroscopic surveys will shrink the errors, as recommended in previous
172 community studies, but they will do even more for particle physics. The richness of the data and detailed

173 attention to the reduction of systematic uncertainties will enable many new tests of the behavior of dark
174 energy and general relativity over a wide range of cosmic distance and time scales.

175 The universe appears to have undergone an enormous expansion in the very first moments after the Big
176 Bang. This phenomenon, known as inflation, helps us understand many subsequent details of how the
177 universe developed. Cosmic microwave background (CMB) experiments will probe the physics of inflation
178 with enough sensitivity to test many of our ideas about the birth of the universe.

179 Cosmic Frontier experiments can also help us understand neutrinos. Studies of the CMB, measurements
180 of the distribution and apparent shapes of galaxies, and polar-ice neutrino experiments will provide precise
181 information about neutrinos, including the mass hierarchy, the sum of their masses, and the number of light
182 neutrinos. These experiments provide information complementary to laboratory studies of neutrinos, and
183 both types combine to create a powerful means for detailed neutrino investigations.

184 Finally, the Snowmass process reiterated the unique information we can gain from studies of cosmic particles
185 and the detection of significant numbers of the highest-energy cosmic rays produced in nature. These studies
186 provide a window on proton interactions at energies beyond the reach of the LHC and include the detection
187 of extremely high-energy neutrinos produced in the interactions of cosmic rays with CMB photons, which
188 will enable the study of neutrino interactions at center-of-mass energies up to 100 TeV.

189 For all of these areas, the Snowmass study identified essential technologies and facilities, the advances
190 required in theoretical models, and experiments with great promise. The largest projects are, appropriately
191 and necessarily, international. The U.S. is still the leader in this quickly evolving area, but other regions
192 with intensive interest in this physics are advancing rapidly.

193 **Theoretical Physics:** Progress in science is based on the interplay between theory and experiment, between
194 having an idea about nature and testing that idea in the laboratory. Neither can move forward without the
195 other. The U.S. has been a world leader for many decades in particle theory, and a sustained strong and
196 vibrant program remains essential for the success of U.S. particle physics. Theoretical physicists have been
197 a driving force in both the development and testing of the Standard Model, including the discovery of the
198 Higgs boson. They play crucial roles in formulating the big questions in the field, setting out hypotheses that
199 address them, and proposing experimental strategies to confirm or refute them. At the same time, theorists
200 seek new structures that might provide unanticipated results. In contrast to experiment, theory depends
201 less on facilities and equipment; the success of the U.S. program rests mainly on principal investigators in
202 universities and national labs, working with postdoctoral fellows and graduate students, and collaborating
203 with both U.S. and international researchers.

204 *Enabling Frontiers*

205 **Accelerator Capabilities:** Accelerator-based experiments continue to be the mainstay of transformational
206 physics on both the Energy and Intensity Frontiers. Progress in these frontiers demands advancing the
207 capabilities of accelerator facilities. The U.S. has been a leader in accelerator physics, its critical and
208 supporting technologies, and the innovative design of research facilities. Its continuing leadership is necessary
209 for future discoveries.

210 The LHC incorporates major U.S. contributions, including high-field superconducting magnets that focus the
211 beams into collision. For future Energy Frontier facilities, U.S. laboratories have pioneered the technology
212 of Nb₃Sn magnets that will permit higher fields than possible with the present LHC technology. The U.S.
213 LHC Accelerator Research Program (LARP) gives the U.S. a world-leading ability to develop high-field
214 superconducting accelerator magnets — a capability central to the LHC luminosity upgrade and for a future
215 proton collider with far greater energy than the LHC. A 100 TeV-class hadron collider (VLHC) is within the

216 development reach of existing materials for a tunnel of 100 km or larger. A collider at even higher energy
217 may require new classes of superconducting magnets and novel ways of handling synchrotron radiation.

218 The ILC, as described in its Technical Design Report, is ready to proceed to construction. Its design
219 incorporates U.S. contributions in accelerator theory, damping ring design, superconducting accelerator
220 technology, and beam control and delivery. Concepts for multi-TeV lepton colliders include the CLIC two-
221 beam accelerator, plasma wakefield accelerators driven either by beams or lasers, and a muon collider.

222 Accelerators for proposed Intensity Frontier experiments need to deliver multi-megawatt proton beams
223 with flexible, experiment-dependent timing structures — demands beyond the capabilities of any existing
224 accelerator. Multi-megawatt beams are the focus of vigorous R&D for both the Project X superconducting
225 linac and DAE δ ALUS cyclotrons. A broad range of highly flexible timing structures is being pursued within
226 Project X.

227 Managing huge stored energy and controlling beam losses to minuscule levels will be essential to operation
228 of all frontier physics accelerators. Specific challenges include generating high-quality beams, modeling
229 beam dynamics, and managing material damage in high-power targets. Maximizing the potential and
230 minimizing the risks in future accelerator projects requires integrated multi-laboratory programs to increase
231 the readiness of critical technologies. Yet engineering-intensive programs must not squeeze out visionary,
232 innovative research in basic accelerator science.

233 A strong U.S. research program of accelerator stewardship benefits all areas of science and industry that use
234 accelerator technology. The broad application of superconducting technology to accelerator-based science
235 exemplifies the payoff of long-range investment that transcends individual projects. Investment in national
236 laboratories and research universities, with their broad expertise and technical infrastructure, will yield new
237 generations of accelerators capable of higher energies, more intense beams, and more efficient operation.

238 **Underground Laboratory Capabilities:** Many experiments searching for dark matter, proton decay,
239 or seeking to determine the properties of neutrinos must be located underground to shield the sensitive
240 experiments from cosmic ray backgrounds. Underground facilities are located or proposed in North and
241 South America, Europe, Asia and in the Antarctic ice. The scope of underground capabilities in all regions
242 is expected to increase by the end of the decade to accommodate the experimental demand. Locating LBNE
243 underground would allow this experiment to realize its full scientific potential and could make it an anchor
244 for a future domestic underground laboratory, bringing to the U.S. a broad range of compelling experiments
245 and researchers from around the world.

246 **Instrumentation:** Instrumentation enables experiments to answer the science questions facing particle
247 physics. Particle physics has a long and distinguished history of inventing, designing, and building the
248 specialized instrumentation required for its experimental research. The physics requirements of many
249 experiments in the Energy, Intensity and Cosmic Frontiers entail very large-scale detectors, but the cost
250 involved in simply scaling up existing technologies is becoming prohibitive. In order for the field to stay
251 competitive, new innovative technologies will need to be developed. Investment in the development of these
252 new technologies is not a luxury but a prerequisite for the sustainability of the field. It has therefore become
253 very important for the particle physics community to establish a mechanism for developing and implementing
254 a coherent vision for the future direction for particle physics instrumentation.

255 The Snowmass study has formulated a vision for a U.S. instrumentation program for particle physics. We
256 identified key barriers to answering the science questions and recognized select technologies for further
257 investment enabling the U.S. to retain a leadership position in a broad global experimental program.

258 Accelerator experiments will require fast, radiation-hard, low-mass, highly segmented, intelligent and so-
259 phisticated trackers and vertex detectors; cost-effective, highly segmented, radiation-hard calorimeters with

260 excellent energy resolution; and high-speed data acquisition systems. Experiments studying particle inter-
261 actions with small cross sections or rare decays typically need cost-effective large volume detectors with
262 efficient background rejection using large area charged particle detection and/or optical readout systems.
263 Some of these experiments need materials and sensors with ultra-low intrinsic radioactivity while others need
264 high-granularity detectors with fast timing and good energy resolution. Many experiments in the cosmic
265 frontier utilize large arrays of ultra-low-noise electromagnetic radiation detectors.

266 Executing a strategy to develop innovative technologies requires integrating the diverse capabilities and
267 resources of universities, national laboratories, other branches of science, and industry into detector R&D
268 collaborations, emphasizing the importance of innovation through a domestic instrumentation development
269 program. A coordinating panel for advanced detectors would help articulate the mission of this program and
270 facilitate its implementation. The goals of the program are to develop both incremental and transformational
271 cost-effective technologies with maximal scientific reach, based on the technological strengths in the U.S. A
272 stable and adequately funded generic instrumentation program will ensure that particle physics invests in
273 its future and establishes a foundation for a competitive, healthy long-term program.

274 **Computing:** Computing is essential to all particle physics experiments and in many areas of theoretical
275 physics. Although some hardware is customized, most of it comes from commercial vendors. Thus, selecting
276 the right hardware and using it efficiently are essential to achieve our scientific goals.

277 We have to train the personnel required to develop and support the parallel programs needed now and in
278 the future. Increased parallelism is required because of changes in chip technology and the evolution of
279 high-performance systems to include multi-core chips and accelerators. It is also important to support the
280 development and implementation of new algorithms in several theoretical areas.

281 Particle physicists should help in the planning of U.S. and international networks, as LHC upgrades will place
282 more demands on the distributed computing systems for ATLAS and CMS. Failure to do so could lead to
283 bottlenecks from wide-area networks, hampering the analysis of results from those and future experiments.
284 Funding agencies should encourage enhanced coordination of software development across all frontiers. The
285 needs of Intensity Frontier experiments are not yet at the level of LHC experiments, but will grow.

286 Future experiments will demand better storage capacity and bandwidth. Disk prices may not drop as rapidly
287 as they have in the past. Scientists working on the cutting edge of computing must also continue to detail
288 their needs to the national supercomputing centers. The funding agencies should seek community input on
289 the appropriate mix of resources dedicated to a particular project and shared computing resources.

290 Early attention to these issues can increase efficiency, reduce costs, enable significantly more realistic
291 theoretical calculations, and avoid computing bottlenecks that could limit scientific progress.

292 *Communication, Education, and Outreach*

293 The particle physics community recognizes the critical importance of consistent and coherent communication,
294 education, and public outreach (CE&O). These foster nationwide support for the field and develop the next
295 generation of scientists, engineers, and scientifically literate citizens. More of us should engage in CE&O
296 activities to translate the American public's fascination with particle physics research into the support
297 necessary to enable the field to answer its biggest questions. Existing activities should be augmented with
298 dedicated personnel who will enhance these efforts, provide nationwide coordination, and spearhead new
299 initiatives. These personnel would develop materials for use in informing the public about direct and indirect
300 applications of particle physics research. They would put in place sustainable methods to collect statistics on
301 workforce development and technology transfer. They could provide professional development opportunities
302 for educators, as well as creating new learning opportunities for students of all ages.

303 *Conclusion*

304 With the completion of the Standard Model, particle physicists now turn their attention to still deeper
305 questions about the nature of matter and the constituents of the universe. This report proposes an ambitious
306 array of new experiments. We consider it realistic to carry out these experiments through a long-term plan
307 and through global partnerships. Particle physicists have been the pioneers of large-scale scientific projects.
308 We have constructed facilities of unprecedented scale, including the Tevatron and the Large Hadron Collider,
309 through decades-long programs requiring world-wide collaboration. These led to discoveries that are the
310 foundation of our current success.

311 Several strategic goals have emerged from the Snowmass study.

- 312 • Probe the highest possible energies and distance scales with the existing and upgraded Large Hadron
313 Collider and reach for even higher precision with a lepton collider; study the properties of the Higgs
314 boson in full detail.
- 315 • Develop technologies for the long-term future to build multi-TeV lepton colliders and 100 TeV hadron
316 colliders.
- 317 • Execute a program with the U.S. as host that provides precision tests of the neutrino sector with an
318 underground detector; search for new physics in quark and lepton decays in conjunction with precision
319 measurements of electric dipole and anomalous magnetic moments.
- 320 • Identify the particles that make up dark matter through complementary experiments deep under-
321 ground, on the Earth's surface, and in space, and determine the properties of the dark sector.
- 322 • Map the evolution of the universe to reveal the origin of cosmic inflation, unravel the mystery of dark
323 energy, and determine the ultimate fate of the cosmos.
- 324 • Invest in the development of new, enabling instrumentation and accelerator technology.
- 325 • Invest in advanced computing technology and programming expertise essential to both experiment and
326 theory.
- 327 • Carry on theoretical work in support of experimental projects and to explore new unifying frameworks.
- 328 • Invest in the training of physicists to develop the most creative minds to generate new ideas in theory
329 and experiment that advance science and benefit the broader society.
- 330 • Increase our efforts to convey the excitement of our field to others.

331 In pursuit of these projects, we have developed a community that links together scientists from all regions
332 of the world pursuing common goals. Our community is ready and eager to carry out the next steps in
333 humankind's quest to understand the basic workings of the universe.

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Summary of the 2013 Community Summer Study

1.1 Introduction

The 2013 Community Summer Study, known as “Snowmass,” sought to identify the critical research directions for the United States particle physics program. This meeting was the culmination of intense work over the past year to define the most important questions for this field and identify the most promising opportunities to address them. The resulting study report, presented here, is a key resource for setting priorities in particle physics.

Through the previous six decades of precision and discovery-level particle physics, we have learned much about the basic laws that govern the Universe. We have uncovered the laws that describe the subnuclear forces and, with the discovery of the Higgs boson, the agent that we believe should give mass to all elementary particles. However, there is still much that we do not understand. The advances in our knowledge of elementary particle physics have sharpened the questions in that domain. Recent discoveries about the matter and energy content of the Universe have brought new questions that are equally fundamental, and equally vexing.

One of the goals of Snowmass was to develop a framework of scientific questions that can form the basis for a future program in particle physics, and to survey experiments that would address these questions. To introduce a summary of the results of Snowmass, we propose a basic set of questions about particle physics whose answers are not yet known. The search for these answers will be carried out with a broad range of experimental methods, cutting across the frontiers around which the Snowmass study was organized.

1. How do we understand the Higgs boson? What principle determines its couplings to quarks and leptons? Why does it condense and acquire a vacuum value throughout the Universe? Is there one Higgs particle or many? Is the Higgs particle elementary or composite?
2. What principle determines the masses and mixings of quarks and leptons? Why is the mixing pattern apparently different for quarks and leptons? Why is there CP violation in quark mixing? Do leptons violate CP?
3. Why are neutrinos so light compared to other matter particles? Are neutrinos their own antiparticles? Are their small masses connected to the presence of a very high mass scale? Are there new interactions that are invisible except through their role in neutrino physics?
4. What mechanism produced the excess of matter over anti-matter that we see in the Universe? Why are the interactions of particles and antiparticles not exactly mirror opposites?
5. Dark matter is the dominant component of mass in the Universe. What is the dark matter made of? Is it composed of one type of new particle or several? What principle determined the current density of dark matter in the Universe? Are the dark matter particles connected to the particles of the Standard Model, or are they part of an entirely new dark sector of particles?

- 379 6. What is dark energy? Is it a static energy per unit volume of the vacuum, or is it dynamical and
380 evolving with the Universe? What principle determines its value?
- 381 7. What did the Universe look like in its earliest moments, and how did it evolve to contain the structures
382 we observe today? The inflationary Universe model requires new fields active in the early Universe.
383 Where did these come from, and how can we probe them today?
- 384 8. Are there additional forces that we have not yet observed? Are there additional quantum numbers
385 associated with new fundamental symmetries? Are the four known forces unified at very short
386 distances? What principles are involved in this unification?
- 387 9. Are there new particles at the TeV energy scale? Such particles are motivated by the problem of the
388 Higgs boson, and by ideas about space-time symmetry such as supersymmetry and extra dimensions.
389 If they exist, how do they acquire mass, and what is their mass spectrum? Do they provide new sources
390 of quark and lepton mixing and CP violation?
- 391 10. Are there new particles that are light and extremely weakly interacting? Such particles are motivated
392 by many issues, including the strong CP problem, dark matter, dark energy, inflation, and attempts
393 to unify the microscopic forces with gravity. What experiments can be used to find evidence for these
394 particles?
- 395 11. Are there extremely massive particles to which we can only couple indirectly at currently accessible
396 energies? Examples of such particles are seesaw heavy neutrinos or Grand Unified scale particles
397 mediating proton decay. How can we demonstrate that these particles exist?

398 The search for answers to these questions is intimately tied to the development of technology. Particle
399 physics experiments and accelerators put extraordinary demands on sensors, precision engineering, and data
400 management, incorporated into devices of very large scale. Our community invents new technologies to
401 address these needs and develops them into usable form. The progress of our field requires both technology
402 development directed at the problems of specific experiments and the development of new technologies
403 that provide higher performance or decreased cost for devices with broad application. This technology
404 development for accelerators and detectors ultimately benefits all of physical science.

405 In many areas of physics experimentation, there are specific technological developments that would be of
406 enormous benefit. Existing technologies are unlikely to meet the science needs of future particle physics
407 experiments. New technologies need to be explored that could lead to transformative advances, enabling
408 cost-effective particle physics experiments but also new initiatives of broad importance.

409 In developing such technologies, we need to address several questions.

- 410 1. Experiments continue to reach for rarer processes, more precise measurements, higher energies and
411 luminosities, and more inclusive observations. How do we achieve the finer granularity, larger volume,
412 more radiation-hard, lower-cost, and higher-speed detectors that will in large part determine our
413 experimental reach?
- 414 2. Paradigm-altering technology developments are occurring in electronics and materials design, poten-
415 tially offering breakthrough capabilities. How can these advances be incorporated into new detectors
416 with improved overall performance? How do we make best use of the resources available in universities,
417 national laboratories, and industry to develop new detector systems?
- 418 3. What technologies will be needed to acquire, analyze, and store the enormous amounts of data from
419 future experiments? Can local intelligence be incorporated to manage data flow? How will we fully
420 and efficiently utilize data stored in large databases?

- 421 4. Scaling of current accelerator designs to higher energy leads to machines of very large size, cost, and
422 power demand. Can new technologies lead to more practical strategies? Is there an ultimate highest
423 energy for colliders?
- 424 5. Proposed experiments at a range of energy scales call for particle beams of extreme brightness. Are
425 there technologies to achieve high beam power in a better-controlled and more cost-effective way?

426 It is important for particle physicists to share the excitement and benefits of our field with a broader public.
427 To that end:

- 428 1. How do we engage particle physicists in communication, education and outreach activities so as to
429 convince policy makers and the public that particle physics is exciting and worth supporting?
- 430 2. How do we educate a talented and diverse group of students who choose to enter particle physics and
431 other STEM careers, including science teaching?

432 In the following chapters, we discuss these issues in more detail and explain how they will be addressed in
433 the coming decades by new initiatives in particle physics. The discussion is organized along the lines of the
434 physics frontiers. The next sections contain the summary of each main program element and the conclusions
435 for each of the frontiers.

436 1.2 Intensity Frontier

437 All frontiers of particle physics aim to discover and understand the constituents of matter and their interac-
438 tions at the highest energies, at the shortest distances, and at the earliest times in the Universe. The Standard
439 Model (SM) fails to explain all observed phenomena: new interactions and yet unseen particles must exist.
440 They may manifest themselves either directly, as new particles, or by causing SM reactions to differ from
441 often very precise predictions. The Intensity Frontier explores these fundamental questions by searching for
442 new physics in extremely rare processes or those forbidden in the SM. This requires the greatest possible
443 beam intensities, as well as massive, ultra-sensitive detectors. Many of these experiments are sensitive to new
444 physics at higher mass scales or weaker interaction strengths than those directly accessible at high-energy
445 colliders, thus providing opportunities for paradigm-changing new discoveries complementary to Energy and
446 Cosmic Frontier experiments.

447 The range of experiments encompassing the Intensity Frontier is broad and diverse. Intense beams of
448 neutrinos aimed over long distances at very large detectors will allow us to explore the neutrino mass
449 hierarchy, search for CP violation and non-standard interactions. The very large detectors, if located
450 underground, will provide increased sensitivity to proton decay. Multi-ton detectors searching for neutrinoless
451 double-beta decay will determine whether neutrinos are their own antiparticles. Intense beams of electrons
452 will enable searches for hidden-sector particles that may mediate dark matter interactions. Extremely rare
453 muon and tau decays, if seen, will signal violation of charged lepton quantum numbers. Measurements of
454 intrinsic lepton properties, such as electric and magnetic dipole moments, are another promising thrust.
455 Rare and CP-violating decays of bottom, charm, and strange particles, measured with unprecedented
456 precision, will clarify the new physics underlying discoveries at the Large Hadron Collider (LHC). In any new
457 physics scenario, Intensity Frontier experiments with sensitivities to very high mass scales will be crucial for
458 exploration.

459 At Snowmass, the Intensity Frontier program was defined in terms of six areas that formed the basis of
460 working groups, with experiments that probe neutrinos, baryon number violation, charged leptons, quark
461 flavor physics, nucleons, nuclei, and atoms, and new light, weakly-coupled particles.

462 The working group reports provide a clear overview of the science program within each area of the Intensity
463 Frontier. They present discovery opportunities for facilities that will be available this decade or will come
464 online during the next decade. Here, we briefly summarize the findings from each working group.

465 **Neutrinos:** Decades of experimental and observational scrutiny have revealed less than a handful of
466 phenomena outside the SM. These include the dark energy and dark matter puzzles, and the existence
467 of non-zero neutrino masses. Neutrino masses represent one of the few experimental pointers towards a new
468 underlying theory. While many experiments continue to look for other new phenomena and deviations from
469 SM predictions, it is clear that continued detailed study of the neutrino sector is of the utmost importance.

470 Compared to the other fermions, the elusive neutrinos have been extremely difficult to study in detail.
471 Despite the challenges, neutrino physics has been tremendously successful over the past two decades. From
472 almost complete lack of knowledge about neutrino mass and mixing twenty years ago, we now have a robust,
473 simple, three-flavor paradigm describing most of the data.

474 However, key questions in the three-flavor sector remain: We do not know the mass ordering or the value of
475 neutrino masses, nor whether neutrinos violate CP symmetry, nor whether the neutrino is its own antiparticle,
476 and we have only just begun to test the three-flavor paradigm. A precision neutrino oscillation program
477 is required to carry out such measurements. Furthermore, some experiments have uncovered intriguing
478 anomalies that merit additional study, and could lead to the discovery of states or interactions beyond
479 the SM. Advances in detector technology and analytical techniques for the next generation of neutrino
480 experiments are well underway. We have clear experimental paths forward for building on our success, for
481 precision testing of the three-flavor paradigm, for the exploration of anomalies, and for the measurement of
482 fundamental neutrino properties and interactions.

483 The next decade promises significant experimental progress around the world. In the search for neutrinoless
484 double-beta decay, a number of experiments rely on complementary isotopes and experimental techniques.
485 The next generation of 100-kg-class neutrinoless double-beta-decay search experiments should have sensitivity
486 to effective masses in the 100 meV range; beyond that, there are opportunities for ton-class experiments
487 that will reach sub-10 meV effective mass sensitivity, pushing below the inverted hierarchy region. The next
488 generation of tritium-beta-decay experiments will directly probe neutrino masses a factor of 10 smaller than
489 the best current bounds. Innovative ideas may help to go beyond these sensitivities.

490 The neutrino mass hierarchy can be unambiguously resolved using accelerator neutrino oscillation experiments
491 with baselines around 1000 km (or longer) and detector masses of order tens of kilotons. Precision measure-
492 ments of atmospheric neutrino oscillations with megaton-scale underground detectors can also resolve the
493 mass hierarchy. The discovery of a non-zero θ_{13} mixing angle enables long-baseline neutrino experiments to
494 search for leptonic CP violation in appearance experiments. The search for CP violation in the neutrino sector
495 is a top priority for particle physics efforts worldwide, and vigorous planning for the next-generation large-
496 scale neutrino oscillation experiment is underway internationally. Regardless of the experimental approach,
497 high-power proton beams (greater than 1 MW) coupled with massive detectors (of order 100 kiloton), are
498 needed to study CP violation in neutrino oscillations. The U.S., with the Long-Baseline Neutrino Experiment
499 (LBNE) and a future multi-megawatt beam from Project X at Fermilab, is uniquely positioned to lead an
500 international campaign to measure CP violation and push the limits of the three-flavor paradigm. An
501 underground location for a far detector significantly enhances the physics reach. LBNE represents a vibrant
502 global program with the U.S. as host.

503 Given the challenges associated with precision measurements in the neutrino sector, complementary baselines,
504 sources, and detector techniques will be required to bring the picture into focus. New accelerator technologies,
505 such as neutrino factories and cyclotron-based-sources, may eventually take measurements to the next level.
506 Smaller experiments will also play a key role in addressing some of the remaining anomalies and hints for
507 physics beyond the three-neutrino paradigm, and study neutrino–matter interactions in detail.

508 The diversity of physics topics that can be probed through the neutrino sector is very significant, and
509 the interplay between neutrino physics and other fields is rich. Neutrinos can and will provide important
510 information on structure formation in the early Universe; Earth, Sun, and supernova physics, nuclear
511 properties; and rare decays of charged leptons and hadrons. The neutrino sector sits at the nexus of a
512 worldwide effort that crosses the frontiers of particle physics.

513 **Baryon number violation:** Within the SM, protons are stable, as baryon number is assumed to be
514 conserved. However, baryon number is not a fundamental symmetry of the SM and is not conserved in many
515 of its extensions. In particular, baryon number violation is an essential ingredient for the creation of the
516 observed asymmetry of matter over anti-matter in the Universe. Grand Unified Theories (GUTs) predict
517 that the proton decays with a lifetime in excess of 10^{30} years, with the decay being mediated at scales of order
518 10^{16} GeV. Two important decay channels in GUTs are $p \rightarrow e^+\pi^0$ and $p \rightarrow \bar{\nu}K^+$, with several other modes
519 also being possible. The current limits on the proton lifetime in these two channels are roughly 10^{34} years
520 and 6×10^{33} years, respectively, which is a factor of 5 to 10 below predictions in certain well-motivated GUT
521 models.

522 The search for proton decay is carried out in detectors containing enough protons, and placed underground
523 to reduce backgrounds. Large neutrino oscillation detectors are ideal for this task, and proton decay is an
524 important piece of their physics portfolio. The largest existing underground neutrino experiment is the 22.5-
525 kiloton Super-Kamiokande water Cherenkov detector. Future underground neutrino experiments, such as a
526 34-kiloton liquid argon TPC (LBNE) or a 560-kiloton water Cherenkov detector (Hyper-Kamiokande) can
527 measure lifetimes on the order of GUT expectations with exposures of roughly 10 years. Typical exposure of
528 these experiments could reach a sensitivity of $\tau(p \rightarrow e^+\pi^0) < 10^{35}$ years and $\tau(p \rightarrow \bar{\nu}K^+) < 3 \times 10^{34}$ years.

529 Neutron-antineutron oscillations would violate baryon number by two units. They are searched for with a
530 beam of free neutrons, in which a neutron would transform into an antineutron that annihilates in a distant
531 detector. Such oscillations are expected in theories where baryogenesis occurs near or below the electroweak
532 scale. A proposed experiment at Project X at Fermilab, using free neutrons from a 1 MW spallation target,
533 could improve existing limits on the oscillation probability by four orders of magnitude.

534 **Charged leptons:** The charged lepton experimental program offers significant discovery opportunities in
535 this decade’s experiments and in even more sensitive experiments possible with future facilities such as
536 Project X at Fermilab. Extremely sensitive searches for rare decays of muons and tau leptons, together with
537 precision measurements of their properties, will elucidate the scale and dynamics of flavor generation or limit
538 the scale of flavor generation to well above 10^4 TeV. Any indication of charged lepton flavor violation (CLFV)
539 would be an indisputable discovery of new physics. Precision measurements of lepton flavor-conserving (LFC)
540 processes can be used to verify predictions of the SM and look for signs of new physics.

541 The experimental program consists of a large and diverse set of opportunities and includes multi-purpose
542 experiments that utilize the large tau production rates at high-luminosity B factories, as well as highly
543 optimized single-purpose experiments that explore muon transitions.

544 Significant advances in studying CLFV in the muon sector can be achieved this decade. For the rare decay
545 $\mu \rightarrow e\gamma$, the MEG upgrade at the Paul Scherrer Institute (PSI) can reach branching fractions up to 6×10^{-14} .
546 The Mu3e collaboration at PSI plans to improve their sensitivity to $\mu \rightarrow 3e$ by approximately four orders of
547 magnitude.

548 Observation of the direct conversion of a muon to an electron in the field of a nucleus would provide a
549 powerful window to physics beyond the SM. Current limits for $\mu N \rightarrow eN$ conversion are at the level of 10^{-12}
550 to 10^{-13} from experiments at PSI. Later this decade, COMET at J-PARC plans to improve these bounds
551 by two orders of magnitude. A separate proposal at J-PARC, DeeMe, would use a different technique to
552 reach a similar sensitivity. Before the end of the decade, Mu2e at Fermilab, followed soon by COMET,
553 will begin operations and improve the existing search reach by four orders of magnitude. This would reach
554 sensitivity to signals from supersymmetric grand unified models. If no signal is observed, this would set
555 constraints on CLFV physics at the scale of 10^4 TeV. Future experiments beyond these are being considered
556 in conjunction with more intense muon beams that could be available with new facilities at J-PARC or
557 Project X at Fermilab.

558 The muon's magnetic moment is predicted very precisely in the SM. New physics contributes via radiative
559 corrections. The present level of sensitivity was obtained by the E821 experiment at Brookhaven National
560 Laboratory, with a difference between the measurement and the SM theoretical prediction of 3.6σ . A new
561 experiment, E989 at Fermilab, will re-use the E821 muon storage ring at Fermilab with the same experimental
562 technique. E989 is expected to increase the statistics by a factor of 20 with a corresponding reduction of
563 systematic uncertainties, resulting in an overall reduction in the experimental error by a factor of roughly
564 4. An alternate approach at J-PARC using lower-energy muons is expected to have the same precision as
565 E989 at Fermilab but very different systematics. A world-wide effort is underway to reduce the theoretical
566 uncertainty in the SM prediction with new data from e^+e^- machines and breakthroughs in lattice gauge
567 theory.

568 Important observables in τ leptons are CLFV decays, CP violation, the electric dipole moment, and the
569 anomalous magnetic dipole moment. New physics effects usually scale as a function of the lepton mass, and
570 hence τ observables can be very sensitive to new contributions. The large τ production rates possible at the
571 future SuperKEKb facility in Japan could achieve an order of magnitude improvement in CLFV branching
572 fractions over current results from BABAR and Belle.

573 The charged lepton sector has significant potential to reveal more information on the fundamental principles
574 of nature, and the U.S. has the opportunity to play a leading role with facilities planned for this decade.

575 **Quark flavor physics:** The study of strange, charm, and bottom quark systems has a long and rich history
576 in particle physics. Measurements of rare processes in the flavor sector have led to startling revelations
577 and played a critical role in the development of the SM. The constraints on physics beyond the SM from
578 flavor physics considerations are powerful. The current quark flavor data set is mostly in agreement with
579 SM expectations with a handful of 3 to 4 *sigma* anomalies. New corrections to the SM at the level of tens
580 of percent are still allowed by the data. Contributions to flavor processes from many theories beyond the
581 SM arise at this level, and thus more precise measurements may observe new physics. If new massive states
582 are observed at the LHC, detailed measurements of the quark flavor sector will be necessary to determine
583 the underlying theory and its flavor structure. If such states are not discovered in high-energy collisions,
584 then precision quark flavor experiments, with their ability to probe mass scales far beyond the reach of the
585 LHC, provide the best opportunity to set the next energy scale to explore. Depending on the strength of
586 new physics interactions, this program already indicates that the new physics scale is above 1 TeV and in
587 some scenarios above 10^5 TeV. Proposed experiments can probe even further. Continued investigations of
588 the quark flavor sector are thus strongly motivated.

589 A well-planned program of flavor physics experiments has the potential to continue this history of advances.
590 Such a program exists worldwide with the LHCb experiment at the LHC, an upgraded SuperKEKB facility
591 in Japan, BESIII in China, and future rare kaon decay experiments at CERN, J-PARC and potentially
592 Fermilab. These facilities will carry out a rich multi-purpose program in the strange, charm, and bottom
593 sectors and perform numerous crucial measurements of rare decays and CP-violating observables. The

594 proposed experiment ORKA at the Fermilab Main Injector would probe rare kaon decays to unprecedented
595 precision and would retain the U.S. capability to perform quark flavor experiments. In the longer term,
596 Project X at Fermilab could become the dominant facility in the world for rare kaon decays. The expected
597 sensitivities for these future programs are detailed in the full report, and are at the level which could discover
598 new physics. It is important to note that these results are not predicated on future theoretical progress,
599 although theoretical advancements will strengthen the program by increasing the set of observables that can
600 reveal new physics. U.S. contributions and support for quark flavor experiments are necessary in order for
601 the U.S. HEP program to have the breadth to assure meaningful participation in future discoveries.

602 **Nucleons, nuclei, and atoms:** The use of nucleons, nuclei, and atoms as laboratories for the study of
603 fundamental interactions is entering a new era. These systems have sensitivity to physics beyond the SM
604 and provide important tests of vital symmetries through measurements of electric dipole moments (EDMs),
605 weak decays of light hadrons, weak neutral currents, and atomic parity violation.

606 Observation of an EDM would signify both parity and time-reversal symmetry violation, and would probe
607 the physics of CP violation. The SM predictions (via multi-loop contributions) for the EDMs of the electron,
608 neutron, and nucleus are 10^{-38} , 10^{-31} , and 10^{-33} e-cm, respectively. EDM measurements are challenging and
609 the present experimental sensitivity is approximately 10^{-27} , 3×10^{-26} , and 3×10^{-29} e-cm for the electron,
610 nucleon, and ^{199}Hg nucleus, respectively. Experiments searching for the electron EDM typically use the
611 polar molecules YbF and ThO and ultimately expect to reach a level of 3×10^{-31} e-cm. Several current or
612 planned experiments searching for the neutron EDM are expected to reach a sensitivity of 5×10^{-28} e-cm, a
613 factor of 100 below current limits. For atoms, future experiments using mercury, radon, and radium expect
614 sensitivities at the level of 10^{-32} e-cm. This would require upgraded facilities such as FRIB at Michigan
615 State University or Project X at Fermilab. These future programs will be sensitive to signals predicted to
616 appear in several theories beyond the SM.

617 The weak decays of light hadrons provide precision input to the SM and are a sensitive test of new interactions.
618 The ratio of decay channels $e\nu/\mu\nu$ for pions and kaons affords a precise test of lepton universality, probing
619 new physics up to 10^3 TeV. For pions, experiments at TRIUMF and PSI will improve the measurement error
620 by a factor of five. Neutron beta decay provides an accurate determination of the CKM element V_{ud} , enabling
621 strong tests of CKM unitarity and constraining new physics up to scales of ~ 10 TeV. Several programs are
622 underway to measure observables of the neutron lifetime and decay asymmetries with improved precision.
623 Measurements of parity-violating asymmetries in fixed-target scattering with polarized electrons allows for
624 a precision determination of the weak mixing angle at low values of momentum transfer, which in turn
625 constrains new parity-violating effects up to $2 - 3$ TeV. An improved polarized Møller scattering experiment
626 at upgraded JLab facilities expects to determine the weak mixing angle to an accuracy comparable to
627 measurements at the Z pole. Parity violation in atomic transitions also yields valuable measurements of the
628 weak mixing angle. New techniques requiring intense sources are being developed; Project X at Fermilab
629 would provide more rare isotopes for this program than any other facility.

630 **New light, weakly coupled particles:** New light particles that couple very weakly to the SM fields are
631 a common feature of extensions beyond the SM. They are motivated by both theoretical and observational
632 considerations, including the strong CP problem and the nature of dark matter and dark energy. Examples
633 of such particles include axions, hidden-sector photons, milli-charged particles, and chameleons. These
634 hidden-sector particles typically couple weakly to the photon via mixing. Intense sources are hence required
635 to produce them at rates sufficient to enable their discovery. The parameters relevant for searches of such
636 hidden-sector particles are their mass and coupling strength to the photon. A variety of experiments constrain
637 part of this parameter space, but much territory is still open for exploration. The current constraints arise
638 from astronomical observations, cosmological arguments, and a variety of laser, heavy-flavor, and fixed-
639 target experiments. Regions that may signal dark matter detection or annihilation and areas that offer an
640 explanation for the present result on the anomalous magnetic moment of the muon have yet to be probed.

641 Numerous laboratory experiments are either in progress or proposed. Two microwave cavity searches for
 642 axions will be underway soon in the U.S., but require further developments to increase their mass reach.
 643 The light-shining-through-walls technique, where photons are injected against an opaque barrier, continues
 644 to explore open regions of parameter space. More advanced technology is needed to make progress in the
 645 mid-term. Axion helioscope searches were first carried out using borrowed magnets, and now require a
 646 custom-built magnet to improve sensitivity. Collider searches for hidden-sector particles can be performed
 647 via the reaction $e^+e^- \rightarrow \gamma\ell^+\ell^-$ at high luminosity e^+e^- factories or in the decays of gauge bosons at
 648 the LHC. Fixed-target experiments using both electron and proton beams are a promising place to search
 649 for hidden-sector particles. The electron beam experiments APEX at the Jefferson National Accelerator
 650 Facility (JLab) and A1 at the University of Mainz have recently performed short test runs and have plans
 651 for more extensive runs this decade. HPS has been approved by JLab and will run after the 12 GeV upgrade.
 652 DarkLight proposes to use the free-electron laser beam at JLab. Proton fixed-target experiments have the
 653 potential to explore regions of parameter space that cannot be probed by any other technique. Neutrino
 654 experiments, such as MiniBooNe, can widen the search to smaller couplings. The intense proton source
 655 at Project X could also provide a powerful extension to the search reach. Impressively large regions of
 656 parameter space are currently unexplored and are ripe for the discovery of light, weakly-coupled particles.

657 **Summary:** The above program exhibits the broad spectrum of science opportunities attainable at the
 658 Intensity Frontier. While each subfield is at a different stage of maturity in terms of testing the SM, the
 659 proposed experiments in each area are poised to have major impact. The programs involving transitions
 660 of heavy quarks, charged leptons, and nucleons, nuclei, and atoms are advanced, with the most precise SM
 661 predictions and a well-developed experimental effort that has spanned decades. In this case, the next level
 662 of experimental precision would reach the point where effects of new TeV-scale interactions are expected to
 663 be observable. More sensitive searches for proton decay and new light, weakly-coupled particles can cover a
 664 large range of parameter space that is consistent with grand unified theories and cosmological observations.
 665 Neutrino physics is just beginning the era of precision measurements where it is possible to probe basic
 666 neutrino properties and answer principal questions. Neutrino physics holds great promise for discovery.

667 Such an extensive program is necessary to address the unresolved fundamental questions about nature. The
 668 knowledge we seek cannot be gained by a single experiment or on a single frontier, but rather from the
 669 combination of results from many distinct approaches working together in concert. The full report from the
 670 Intensity Frontier provides a reference for the captivating science that can be carried out in this decade and
 671 next. The Intensity Frontier program has the potential to make discoveries that change paradigms and alter
 672 our view of the Universe.

673 1.3 Energy Frontier

674 Experiments at the Energy Frontier make use of high-energy accelerators to produce and study heavy
 675 elementary particles and to search for new ones. The Energy Frontier includes experiments at the Large
 676 Hadron Collider at CERN and those at future colliding-beam accelerators proposed for lepton-lepton and
 677 proton-proton collisions.

678 The first run of the LHC has closed a nearly half-century-old chapter in the story of elementary particle
 679 physics. We have discovered a most unusual new particle with properties very similar to those expected of the
 680 Standard Model Higgs boson. The appearance of this particle — and further confirmation of its identity —
 681 ends one era and opens another. On one hand, the Standard Model of particle physics is complete. We know
 682 all of the particles in this model and how they interact with one another and we have at least a basic idea
 683 of their properties. On the other hand, we also know that the Standard Model is incomplete in important
 684 ways. It challenges us to uncover the physics behind its apparently ad hoc structure. We are certain that

685 a host of observed anomalous phenomena and set of confusing conceptual questions have explanations that
686 require new physics outside the Standard Model.

687 The LHC and the CMS, ATLAS, and LHCb detectors have brought to bear impressive capabilities for
688 exploring the answers to these new questions. The LHC accelerator is expected to dramatically increase its
689 ability to deliver beams in the period between now and 2030, increasing its energy by almost a factor of two
690 and its integrated luminosity by a factor of 100. The detectors will improve their ability to collect enormous
691 data sets and to discriminate the properties of events with increasing precision. Around the world, other new
692 accelerators are being considered that will give us additional power in understanding the heaviest particles
693 of the Standard Model and exploring for new ones. In this report, and in the detailed working group reports,
694 we trace out the programs of these accelerators and present their most important goals.

695 **Importance of the TeV energy scale:** Our successful theory of weak interactions is based on the idea
696 of an underlying symmetry that is spontaneously broken. The symmetry of the theory of weak interactions
697 dictates the couplings of the quarks and leptons to the W and Z bosons. Its predictions have been confirmed
698 by high-precision experiments. However, this symmetry forbids the quarks, leptons, and vector bosons from
699 having mass. To reconcile the symmetry of weak interactions with the reality of particle masses, one more
700 unexpected element is required. This is a field or set of fields that couples to all types of particles and forms
701 a condensate filling the Universe. The discovery of the Higgs particle establishes that this condensate exists
702 and is the origin of particle masses.

703 This is an historic achievement. It is not an end but a beginning. It highlights many questions that the
704 Standard Model leaves unanswered. These require new, equally bold ideas. Two of these questions — the
705 nature of the Higgs field and the composition of dark matter — give particularly strong motivations for
706 collider experiments.

707 The Standard Model does not explain the underlying structure of the Higgs field or the reason why it
708 condenses. It does not explain the magnitude of the condensate, which sets the mass scale of all known
709 elementary particles. The fact that the observed Higgs particle is a scalar particle makes it very difficult to
710 understand why this scale is smaller than other basic mass scales of nature such as the Planck scale. There
711 are no simple models that answer this question. New fundamental structures are needed. The Higgs field
712 must be a composite of more basic entities, or space-time itself must be extended, through supersymmetry
713 or through extra dimensions of space. These ideas predict a rich spectrum of new elementary particles,
714 typically including a larger set of Higgs bosons, with masses at the TeV energy scale.

715 The Standard Model also does not account for the dark matter that makes up most of the matter of the
716 Universe. The simplest and most compelling model of dark matter is that it is composed of a stable, weakly
717 interacting, massive particle (WIMP) that was produced in the hot early Universe. To obtain the observed
718 density of dark matter, this model requires the WIMP interactions to be roughly at the TeV energy scale.
719 If this model is correct, it may be possible to study dark matter under controlled laboratory conditions in
720 collider experiments.

721 *Compelling ideas about fundamental physics predict new particles at the TeV energy scale that should be*
722 *discoverable in experiments at the LHC and planned future accelerators. These experiments will provide the*
723 *crucial tests of those ideas. Furthermore, if such particles are discovered, they can be studied in detail in*
724 *collider experiments to determine their properties and to establish new fundamental laws of nature.*

725 The past successes of particle physics and its current central questions then call for a three-pronged program
726 of research in collider experiments:

- 727 1. We must study the Higgs boson itself in as much detail as possible, searching for signs of a larger Higgs
728 sector and the effects of new heavy particles.

- 729 2. We must search for the imprint of the Higgs boson and its possible partners on the couplings of the W
730 and Z bosons and the top quark.
- 731 3. We must search directly for new particles with TeV masses that can address important problems in
732 fundamental physics.

733 The Energy Frontier study pointed to all three of these approaches as motivations for further experiments
734 at colliders. The results of the study confirmed that the existing LHC detectors and their planned upgrades,
735 together with proposed precision lepton collider experiments, will be nimble and sensitive enough to carry
736 this three-fold campaign forward into the next two decades.

737 The Energy Frontier study was organized into six working groups — on the Higgs boson, the W and Z
738 bosons, Quantum Chromodynamics (QCD), the top quark, new particles and forces, and flavor interactions
739 at high energies. Each working group was asked to evaluate the future program for its topic both from a
740 high-level perspective and from the viewpoint of supplying motivation for experiments at a range of proposed
741 accelerators. In the remainder of this section, we present the conclusions of these reports, first by physics
742 topic, then by facility.

743 **Higgs boson:** A new bosonic resonance at 125 GeV was discovered at the LHC only one year ago. Many
744 properties of this particle have now been measured and, up to this point, are consistent with those of the
745 Higgs boson of the minimal SM. The couplings of this boson roughly scale with mass. The specific form of
746 the coupling to the Z boson indicates that the particle has spin-parity 0^+ and that the corresponding field
747 has a nonzero vacuum expectation value.

748 However, we cannot be complacent about the identity and role of this particle. On one hand, the idea that
749 a single scalar field is solely responsible for the generation of all particle masses is just one possibility among
750 many and needs explicit verification. On the other hand, models with additional Higgs bosons and related
751 new particles, and models in which the Higgs boson is composite, are hardly tested. Deviations from the
752 minimal Higgs boson properties due to new particles with mass M are suppressed by a factor $(m_h/M)^2$, so
753 to the extent that the LHC has set lower limits on the masses of new particles at many hundreds of GeV,
754 we would not yet have expected to see the modifications to the Higgs properties caused by those particles.

755 An experimental program to probe the Higgs boson contains several elements. The first is to search for
756 deviations from the minimal SM expectation that the Higgs boson couples to each particle species according
757 to its mass. Such deviations are expected in almost all models of new physics. However, the effects are
758 expected to be small, at the few-percent level if induced by new particles that will not be directly detected at
759 the LHC. There is a characteristic pattern of deviations for each new physics model. The High-Luminosity
760 LHC (HL-LHC) is expected to measure these couplings with precisions of several percent, varying from
761 coupling to coupling. Lepton collider experiments have the potential to push these precisions to the sub-
762 percent level, which would be needed to uncover deviations from Standard Model predictions with significance
763 high enough to claim evidence of new physics.

764 Such a program of precision measurements of Higgs couplings requires a parallel concerted effort in precision
765 theory. It also requires improvement of our knowledge of crucial input parameters such as α_s and m_b , which
766 can be provided by lattice gauge theory computations. Collider experiments can also probe the nonlinear
767 Higgs field self-coupling to the 10–20% level, thereby testing the critically important question of the shape
768 of the Higgs potential.

769 Future experiments should also improve our knowledge of the Higgs boson mass and quantum numbers.
770 The spin of the observed resonance should already be clear from LHC data in this decade. A more subtle
771 question is whether this particle contains a small admixture of a CP-odd state, signaling CP violation in the

772 Higgs sector and confirmation of at least one additional Higgs-like particle. We discuss probes for this effect
773 at various colliders.

774 Finally, it is important to search directly for additional Higgs bosons. The LHC can probe to masses of
775 1 TeV with model-dependent limits. Lepton colliders can make more model-independent searches to
776 masses close to the collider beam energy.

777 ***W* and *Z* boson, QCD, and the top quark:** The study of *W* and *Z* bosons includes both the extension
778 of the program of precision electroweak measurements, and the search for new interactions in the three- and
779 four-vector boson couplings.

780 The minimal SM makes precise predictions for the well-studied precision observables M_W and $\sin^2 \theta_W$. At the
781 moment, the observed values are within 2σ of the predictions; the deviations are consistent with the effects
782 of new particles in a range of new physics models. Better precision in this program is clearly needed. Future
783 experiments will sharpen our knowledge of these quantities and potentially expose inconsistency with the
784 SM. The LHC, especially in its high-luminosity phase, has the potential to reduce the error on the *W* mass to
785 ± 5 MeV. This requires a factor of 7 decrease in the current error due to parton distribution functions and is
786 a challenge to QCD researchers. Lepton colliders can make further improvements, to an error of ± 2.5 MeV,
787 with a dedicated measurement of the *WW* threshold. A linear collider with beam polarization running at
788 the *Z* resonance to produce 10^9 *Z* bosons (Giga-*Z*) is expected to reduce the error on $\sin^2 \theta_W$ by a factor of
789 10. Finally, a circular e^+e^- collider operating in a 100 km tunnel can potentially push both errors down by
790 another factor of 4. All of these precision measurements challenge the inflexible correlations among the SM
791 particles and their respective forces. Such precision measurements of electroweak observables could become
792 discoveries of new physics if the tight constraints within the SM begin to unravel.

793 The second theme of *W* and *Z* boson studies is the search for anomalous nonlinear couplings of the vector
794 bosons. Collider experiments with enough energy to produce pairs of *W* and *Z* bosons are sensitive to three-
795 gauge-boson couplings. At the LHC, we will be sensitive, for the first time, to non-standard four-boson
796 interactions, which would indicate new interactions in vector boson scattering. Lepton collider experiments
797 have the potential to push current uncertainties on three-boson couplings down by an order of magnitude,
798 into the region in which new physics effects are predicted in models in which the Higgs boson is composite.
799 Both hadron and lepton colliders can access vector boson scattering, but the total center-of-mass energy
800 available in a scattering process is a crucial factor. The high-luminosity LHC will be sensitive to vector
801 boson or Higgs resonances with masses well above 1 TeV.

802 QCD is well established as the correct theory of the strong interactions. Nevertheless, advances in QCD
803 are needed to achieve the goals of future experiments, especially at hadron colliders. These experiments
804 require improved knowledge of the parton distribution functions. That can be achieved with data expected
805 from the LHC on the rapidity distributions of *W*, *Z*, and top quark production. In addition, precision
806 cross-section computations, to the NNLO level, are needed for many two- and three-particle production
807 processes, especially those involving the Higgs boson. This will require advances in the theoretical art of
808 QCD computation. Finally, it is important to push the error on the value of α_s below the 0.5% level. Lattice
809 gauge theory seems to be a promising avenue for achieving this goal.

810 The top quark was discovered at the Fermilab Tevatron and studied there with samples of tens of thousands
811 of $t\bar{t}$ pairs. The LHC experiments will produce and study billions of top quarks. At future lepton colliders,
812 we will use the electroweak couplings of top quarks as a production mode and probe these with polarization
813 observables. Both methods will transform our knowledge of this quark, whose properties are intimately
814 connected to the mysteries of flavor and mass generation. To this day, we are surprised at the high mass of
815 this presumably fundamental particle and its proximity to the value of the Higgs vacuum expectation value.

816 The top quark mass is not only an important puzzle in itself but also is an important input parameter
817 for particle physics. The strongest demands on precision in the top quark mass come from the precision
818 electroweak program, where interpretation of a 5 MeV error in m_W requires a 500 MeV error on m_t . This
819 mass must be a theoretically well-defined quantity, convertible to a short-distance parameter such as the \overline{MS}
820 mass. There are strategies applicable at the LHC that allow the measurement of a well-defined top quark
821 mass to this 500 MeV accuracy. At lepton colliders, measurement of the cross section at the top quark pair
822 production threshold gives the \overline{MS} mass to 100 MeV, as required for the more accurate precision electroweak
823 program available at these machines.

824 Top quark couplings will be studied with high accuracy both at hadron and at lepton colliders. New physics
825 from top quark and Higgs compositeness can create few-percent corrections to the gluon, photon, and,
826 especially, Z boson couplings. These effects can be observed as corrections to the pair-production cross
827 sections relative to the predictions of the SM. The top-quark coupling capabilities of a lepton collider are
828 especially strong, with accuracies possible at the sub-percent level. The billions of top quarks produced
829 at the high-luminosity LHC allow very sensitive studies of rare flavor-changing top decays, to a level that
830 complements searches at low energy for flavor-changing quark decays.

831 Models of the Higgs potential and its symmetry breaking typically require new particles that are partners,
832 in some way, of the top quark. The LHC, especially in its high-luminosity stage, will have the capability for
833 extensive searches for supersymmetric partners of the top quark, heavy vector-like top quarks that appear
834 in models with Higgs and top quark compositeness, and heavy resonances that decay to $t\bar{t}$, which appear in
835 models with new space dimensions.

836 **Searches for new particles and interactions:** High-energy colliders can search for new particles with a
837 very broad range of properties. These particles, with masses near the 1 TeV scale, are required in models
838 of electroweak symmetry breaking. Other questions also call for new particles accessible to high energy
839 colliders. A large class of models of dark matter place the dark matter particle as the lightest particle of
840 a TeV mass spectroscopy. Grand unification requires new particles near the TeV scale, including partners
841 of known particles and perhaps also new vector bosons associated with enhanced gauge symmetry. CP
842 violation in the Higgs boson sector is required in models that generate the matter-antimatter asymmetry
843 at the electroweak phase transition. More generally, new particles can bring new sources of flavor and CP
844 violation that might be reflected in the discovery of new flavor-changing reactions at low energy.

845 The LHC has already, in only its first run, increased the reach and power of searches for new particles
846 over a broad scope. We expect that this power will increase dramatically in the next decade, as the LHC
847 experiments acquire 300 fb^{-1} of data at 14 TeV. This extension probes deeply into the region expected for
848 the masses of new particle in all classes of models of electroweak symmetry breaking. The high-luminosity
849 stage of the LHC, up to 3000 fb^{-1} , will provide a further, very significant, extension of the search region.
850 This extension is particularly powerful for states produced through electroweak interactions, for which a
851 factor of 2 increase in the mass reach is available in some cases.

852 Lepton colliders would bring new and complementary capabilities. They would carry out model-independent
853 searches for states such as dark matter candidate particles whose signatures are especially difficult to observe
854 at hadron colliders. Lepton colliders would uncover new decay modes and measure branching ratios and
855 quantum numbers for any new particle within their energy range.

856 **Physics opportunities for colliders:** The physics opportunities described above are reflected as motiva-
857 tions for current and future high-energy colliders. Our study considered a wide range of proposed machines.
858 The full report from the Energy Frontier presents the cases for these machines in some detail.

859 We first point out the opportunity provided by the 14 TeV run of the LHC scheduled for the next decade.
860 This will provide robust searches for new particles over a broad front, with great promise of the discovery of

861 the TeV particle spectrum motivated at the beginning of this section. Any plan for high energy physics in
862 the longer term must include the possibility of discovering new particles in this period and exploiting that
863 discovery at the facilities that will follow.

864 We find the case for the high-luminosity stage of the LHC compelling. This plan to deliver 3000 fb^{-1} has
865 been listed in the European Strategy for Particle Physics as the highest priority accelerator project in Europe
866 for the 2020's. We find that it will provide a significant additional step in the search for new particles, and
867 that it will provide other important capabilities. The most important of these is the beginning of the era
868 of precision Higgs boson measurements, to few-percent precision. It is likely to give the first evidence of
869 the Higgs boson self-coupling. It will provide a program of precision measurement in the SM that will
870 dramatically tighten our knowledge of the W boson and the top quark, with measurements sensitive to the
871 predictions of a variety of new physics models. We have already noted that the additional luminosity will
872 significantly enhance the capability of the LHC to search for new heavy particles.

873 We considered the scientific case for the International Linear Collider (ILC). This next-stage lepton collider
874 has recently completed its Technical Design Report and was judged in the Snowmass study to be ready for
875 construction. This facility is named as the highest priority for new initiatives by the Japanese high-energy
876 physics community. We find that this machine is strongly motivated. It will reach sub-percent accuracy in
877 the study of the Higgs boson, allowing discovery of percent-level effects in the Higgs couplings predicted in
878 new physics models. It will measure the Higgs width in a model-independent way. It will give the capability
879 to observe all possible Higgs modes, including decays to SM modes not observable at the LHC, to dark
880 matter, and to other invisible and exotic states. It will extend our knowledge of the top quark and the W
881 and Z bosons well beyond the precision achievable at the LHC, setting up a confrontation with models that
882 include Higgs boson and top quark composite structure.

883 The Energy Frontier study considered many other accelerator facilities for construction over longer time
884 scales. These included higher energy linear colliders, circular e^+e^- colliders, muon colliders, and photon
885 colliders. We present a detailed discussion of the physics motivations for these facilities in our full report.
886 There was particular interest in a proton collider of energy 100 TeV (VLHC), which would come close to
887 the capability of covering the full model space for models of "natural" electroweak symmetry breaking and
888 WIMP dark matter. Our study developed materials and resources to begin a more complete survey of
889 physics at such a high-energy collider. This study, and a parallel development of magnet technology for
890 higher-energy proton colliders, should be pursued over the next decade.

891 **Conclusions:** Previous surveys of the prospects for high energy accelerator experiments have spoken in
892 terms of reducing the space of parameters — couplings, mixings, masses — as if that were the goal. Now,
893 more than ever, the momentum points not toward exclusion, but toward the discovery of new states. Many
894 possible directions are open and must be pursued.

895 The Higgs boson discovery changes everything. It transforms the research agenda for particle physics, giving
896 us a set of sharp questions that we cannot ignore. It motivates more strongly the exploration of the TeV
897 energy scale, where the solution to the mystery of dark matter and other key problems might also be found.
898 The study of the Higgs particle to high precision, together with high-precision studies of the W , Z , and top
899 quark and searches for new states, provide us with complementary routes to fully explore the particles and
900 forces in this range of energies. The current LHC detectors and their planned upgrades are well suited to
901 carry on this program. Future accelerators will bring new capabilities to pursue it further.

902 High-energy colliders provide manifest opportunities to discover new fundamental interactions of broad
903 consequence. U.S. physicists have been leaders in Energy Frontier experiments up to now and are well
904 positioned to take a leading role in the discoveries of the coming decades.

1.4 Cosmic Frontier

Investigations at the Cosmic Frontier use the Universe as a laboratory to learn about particle physics. Our understanding of the Universe has been transformed in recent years. In particular, experiments at the Cosmic Frontier have demonstrated that only 5% of the contents of the Universe are well understood, with the rest composed of mysterious dark matter and dark energy. As a result, the Cosmic Frontier now plays a central role in the global particle physics program, providing overwhelming evidence for new particles and new interactions, as well as powerful, unique opportunities to address many of our most fascinating questions: What is dark matter? What is dark energy? Why is there more matter than antimatter? What are the properties of neutrinos? How did the Universe begin? What is the physics of the Universe at the highest energies?

To identify outstanding scientific opportunities for the coming 10 to 20 years, the Cosmic Frontier Working Group was organized into six subgroups: Weakly-interacting massive particle (WIMP) dark matter direct detection, WIMP dark matter indirect detection, non-WIMP dark matter, dark matter complementarity, dark energy and CMB, and cosmic particles and fundamental physics. In several cases, these subgroups were further divided into topical working groups.

The Λ CDM standard model of cosmology provides the backdrop for much of Cosmic Frontier research. In this model, the Universe underwent a very early epoch of accelerated expansion (inflation), which was followed by eras in which the Universe was dominated successively by radiation, cold dark matter (CDM), and dark energy (Λ). At present, the known particles make up only 5% of the energy density of the Universe, with neutrinos contributing at least 0.1%. The rest is 25% dark matter and 70% dark energy. Remarkably, incisive measurements that explore all of the key components of the model are now within reach. The leaps in sensitivity of the new facilities bring us to a time with strong discovery potential in many areas. Further surprises are likely in this rapidly advancing area, with potentially far-reaching consequences.

Dark matter: The work of Snowmass highlighted the coming decade as one of particular promise for the goal of identifying dark matter. Evidence for particle dark matter has been building for 80 years through the study of galaxy clusters, galactic rotation curves, weak lensing, strong lensing, hot gas in galaxy clusters, galaxy cluster collisions, supernovae, and the cosmic microwave background (CMB). However, all evidence so far is based on dark matter's gravitational interactions, and its particle identity remains a deep mystery.

Among the many dark matter candidates, one well-known possibility is weakly-interacting massive particles with masses in the 1 GeV to 100 TeV range. Particles with these properties appear in many models designed to address the gauge hierarchy problem. In cosmology, particles with these properties may obtain the correct relic density either through thermal freeze-out or through an asymmetry connecting their number density to that of baryons.

WIMP direct detection experiments search for the interactions of WIMPs with normal matter. WIMPs may scatter elastically off nuclei, producing recoil energies in the 1–100 keV range, which can be detected through phonons, ionization, scintillation, or other methods. There are daunting backgrounds, and direct detection experiments must be placed deep underground. In the last several years, however, this field has seen a burgeoning of innovative approaches to discriminate signal from background, including experiments incorporating dual-phase media, self-shielding, pulse shape discrimination, and threshold detectors.

The first two decades of direct detection experiments have yielded a diverse and successful program, resulting in “Moore’s Law”-type progress, with sensitivities doubling roughly every 18 months. In the coming decade, this rate of progress is expected to continue or even accelerate for both spin-independent and spin-dependent interactions. Upcoming second-generation (G2) experiments will improve sensitivities by an order of magnitude, probing the Higgs-mediated cross sections expected for well-known supersymmetric and extra-

949 dimensional candidates, and also extending the sensitivity to both \sim GeV low-mass WIMPs, where possible
950 signals have been reported, and \sim TeV masses that are beyond the reach of colliders. Following these
951 experiments, multi-ton-scale third-generation (G3) experiments are expected to improve current sensitivities
952 by up to three orders of magnitude and will either find dark matter or detect background events from
953 solar, atmospheric, and diffuse supernovae neutrinos. Probing beyond this sensitivity will require either
954 background subtraction or techniques such as directional detection or annual modulation. The Snowmass
955 process produced a detailed census of present and proposed direct detection facilities, with uniform treatment
956 of their capabilities and issues, along with a survey of promising technologies.

957 WIMPs may also be found through indirect detection, in which pairs of WIMPs annihilate, producing SM
958 particles, including gamma rays, neutrinos, electrons and positrons, protons and antiprotons, and deuterons
959 and antideuterons. Detection of these particles may be used to constrain or infer dark matter properties.
960 The expectation that WIMP annihilation in the early Universe determines the dark matter abundance sets
961 a natural velocity-averaged annihilation cross section of $\langle\sigma_{\text{ann}}v\rangle \sim 3 \times 10^{-26} \text{ cm}^3\text{s}^{-1}$ for indirect detection
962 experiments.

963 Gamma rays from dark matter annihilation may be detected by both space- and ground-based experiments.
964 In space, the Fermi-LAT has recently demonstrated the promise of this approach, excluding the natural cross
965 section $\langle\sigma_{\text{ann}}v\rangle$ for dark matter masses below 30 GeV, given certain halo profile and annihilation channel
966 assumptions. The reach is expected to be extended significantly with additional data. On the ground,
967 VERITAS and other atmospheric Cherenkov telescopes have set significant limits by looking for gamma rays
968 from dark matter-rich dwarf galaxies. Moving forward, the atmospheric Cherenkov telescope community has
969 coalesced to build the Cherenkov Telescope Array (CTA), with sensitivity at the natural cross-section scale
970 for dark matter masses from 100 GeV to 10 TeV, far beyond current or planned colliders. These projections
971 require U.S. involvement in CTA, which will double the planned mid-sized telescope array and enable critical
972 improvements in sensitivity and angular resolution.

973 Neutrinos also provide promising means for indirect detection of dark matter. High-energy neutrinos from
974 the core of the Sun would be a smoking-gun signal of dark matter particle annihilation. The signal depends
975 primarily on the spin-dependent WIMP-nucleon scattering cross section, which determines the capture rate.
976 Current bounds from Super-K in Japan and IceCube at the South Pole already provide leading limits on this
977 cross section, and PINGU, an infill array upgrade to IceCube, will extend the sensitivity to lower masses.
978 In the coming decade, IceCube and PINGU, along with Hyper-Kamiokande, will probe cross sections one
979 to two orders of magnitude below current bounds, with sensitivities competitive with those of planned G2
980 direct detection experiments.

981 Antimatter signals of dark matter are pursued in a variety of ways. Recent measurements of cosmic-
982 ray positrons by the AMS-02 magnetic spectrometer confirm and improve with excellent precision earlier
983 measurements by PAMELA and Fermi. The rising positron fraction could be indicative of positrons created
984 in the decay or annihilation of dark matter. In the near future, AMS-02 will extend its determination of
985 the positron fraction to energies close to 1 TeV, and add important information on cosmic-ray propagation.
986 Given the possibility of astrophysical sources of primary positrons, however, it may be very difficult to
987 definitively attribute the excess positrons to dark matter. Antideuterons provide a signal that is potentially
988 more easily discriminated from astrophysical backgrounds. With a long-duration balloon flight, the General
989 Antiparticle Spectrometer (GAPS) detector could provide sensitivities comparable to AMS-02. Last, the
990 production of positrons and electrons from dark matter annihilation also produces secondary radiation.
991 Detection of signals with radio to X-ray frequencies has the potential to probe the WIMP parameter space.

992 The Snowmass process also evaluated the prospects for non-WIMP candidates, which could be some or all of
993 the dark matter. The axion is particularly well-motivated, as it arises from the leading solution to the strong
994 CP problem of the SM. RF-cavity and solar searches for axions, such as ADMX and IAXO, will probe a

995 large range of axion parameter space, including the cosmologically-favored region, and have strong discovery
996 potential. Sterile neutrinos are also highly motivated by the observed non-zero masses of active neutrinos.
997 In the mass range where sterile neutrinos are dark matter candidates, their radiative decays produce a
998 monoenergetic photon, which may be detected with X-ray telescopes. Many other dark matter candidates
999 were also surveyed, including asymmetric dark matter, primordial black holes, Q-balls, self-interacting dark
1000 matter, superheavy dark matter, and superWIMP dark matter.

1001 How do the diverse strategies for identifying dark matter fit together? The Snowmass process produced a
1002 clear articulation of how the different approaches—including the direct and indirect detection experiments
1003 mentioned above, but also particle colliders and astrophysical probes—each provide unique and necessary
1004 information. This complementarity was examined in two theoretical frameworks. First, the discovery
1005 prospects were examined in complete supersymmetric models, with randomly selected parameters in the
1006 phenomenological MSSM framework. Second, the possibility that only the dark matter particle is kinematically
1007 accessible was considered using the framework of dark matter effective theories. In both cases,
1008 the complementarity of different approaches was evident at all levels, both to establish a compelling dark
1009 matter signal and, just as importantly, after discovery, to determine the detailed properties of the particle
1010 or particles that make up dark matter.

1011 **Dark energy and CMB:** Cosmic surveys — optical imaging and spectroscopic surveys and detailed
1012 measurements of the CMB — precisely map the Universe on many different angular scales and over wide
1013 ranges of cosmic time. They provide unique information about cosmology and new physics, including
1014 inflation, dark matter, dark energy, and neutrino properties. These measurements are challenging, requiring
1015 advances in instrumentation and excellent control of systematic effects. Fortunately, these advances are
1016 now within reach, thanks to decades of investment and close collaborations between particle physicists and
1017 astrophysicists. The payoffs for this effort are large.

1018 Measurements of the distance-redshift relation, first using supernovae and then additional complementary
1019 techniques, revealed the expansion history of the Universe, particularly over the past several billion years,
1020 and yielded the surprising discovery that the expansion rate has been increasing instead of decreasing. Now
1021 we must determine what is causing the cosmic acceleration. This “dark energy” must produce negative
1022 pressure to be responsible for the observed effect. One important clue is whether the negative pressure
1023 has been constant in time or is evolving. The stage III (the DES and HSC imaging surveys, and the PFS
1024 and eBOSS spectroscopic surveys) and stage IV (LSST imaging survey and DESI spectroscopic survey on
1025 mountaintops; Euclid and WFIRST-AFTA in space) dark energy facilities will constrain both the value and
1026 the evolution of the value with much higher precision, as recommended in previous community studies, but
1027 they will also do much more. We must also check whether our description of gravity is correct, and this is
1028 where measurements of the growth of structure, over a wide range of distance scales using both imaging and
1029 spectroscopic surveys, are needed.

1030 There are several alternatives to general relativity (GR) that can accurately describe the observed distance-
1031 redshift relation, but they also modify the behavior of gravity over different distance scales. The alternative
1032 models therefore predict structure growth rates that are different from those in the standard theory. Mea-
1033 suring the structure growth rate over many different distance scales will test GR and the alternative models.
1034 Deviation from expectation on just one of these scales will signal new physics. In other words, the upcoming
1035 dark energy facilities, particularly at stage IV, where systematic error management is built deeply into the
1036 design, will provide many precise tests and will characterize the behavior of dark energy beyond merely
1037 a single parameter value and its evolution with time. We will know the strength of the effects in a two-
1038 dimensional parameter space of distance and cosmic time, as well as any deviations from expectations in
1039 their correlations. Further surprises may await us.

1040 Inflation is the leading paradigm for the dynamics of the very early Universe, and current observations of
1041 large-scale structure lend support to this intriguing idea. The most direct available probes of inflation come
1042 from CMB observations, and the overall agreement is remarkably good. However, it has not been possible to
1043 explore the underlying physics of inflation until now: the coming generations of CMB experiments will have
1044 sufficient sensitivity to falsify large classes of models. The signal is a characteristic pattern with non-zero
1045 curl (called “*B* mode”), faintly imprinted on the polarization of the CMB fluctuations, due to gravitational
1046 waves produced during the epoch of inflation. The shape of the potential of the scalar field driving inflation
1047 directly affects the spectrum of gravitational waves and hence the strength of the imprint, r (the ratio of
1048 tensor to scalar power), over characteristic angular scales on the sky. The current generation of experiments
1049 is sensitive to $r \sim 0.1$, but over the next 10 to 20 years, improvements of two orders of magnitude are
1050 possible by scaling the number of detectors by similar factors, from $\sim 10^3$ (current) to $\sim 10^4$ (generation III)
1051 to $\sim 5 \times 10^5$ (generation IV). This would require a change from the way things have been done in the past.
1052 Groups would merge into one coordinated effort, tapping national lab facility design, integration, computing,
1053 and management capabilities.

1054 In addition, future optical and CMB cosmic surveys, as well as future polar-ice neutrino projects (see
1055 below), will provide precise information about neutrino properties, including the mass hierarchy, the number
1056 of light neutrinos, and the sum of the neutrino masses. Combining this with information from accelerator-
1057 and reactor-based neutrino experiments, as well from experiments searching for neutrinoless double-beta
1058 decay, will accelerate our understanding of fundamental neutrino properties and enable us to understand the
1059 implications of apparent inconsistencies.

1060 Snowmass provided an excellent opportunity to address common problems and to develop a common vision
1061 for the potential of cosmic surveys to advance particle physics. Highlights included developing detailed
1062 strategies to distinguish dark energy from modified gravity; exploiting the complementarity of probes
1063 for determining the key cosmological parameters; understanding more deeply the strengths and ultimate
1064 limitations of the different techniques; and discussing the planned facilities, which are the result of intensive
1065 community processes over many years. The group articulated a set of goals: (1) remain a leader in dark
1066 energy research, (2) build a generation IV CMB polarization experiment, and (3) extend the reach of cosmic
1067 surveys with targeted calibration campaigns, targeted R&D, and support for work at the interface of theory,
1068 simulation, and data analysis.

1069 **Cosmic particles:** Measurements of fluxes of cosmic particles (charged particles, photons, and neutrinos)
1070 also address many topics in particle physics beyond indirect dark matter searches. Recent results include
1071 the detection by IceCube of very high-energy neutrinos that are likely to be from astrophysical sources;
1072 the observation of the GZK suppression in the cosmic-ray flux above 3×10^{19} eV; the measurement of the
1073 positron fraction up to 300 GeV, suggesting the existence of primary sources of positrons from astrophysical
1074 processes and/or dark matter interactions; and confirmation that supernova remnant systems are a source
1075 of galactic cosmic rays. These and other discoveries were made by the current generation of experiments.

1076 Goals for the coming decade include determining the origin of the highest energy particles in the Universe,
1077 measuring interaction cross sections at energies unattainable in terrestrial accelerators, detecting the GZK
1078 neutrinos that arise from the interactions of ultra-high-energy cosmic rays with the CMB, determining the
1079 neutrino mass hierarchy, and searching for other physics beyond the Standard Model.

1080 To meet these goals, the group recommends: significant U.S. participation in the Cherenkov Telescope Array
1081 (CTA), which is the next-generation ground-based gamma-ray facility; simultaneous operation of Fermi,
1082 HAWC, and VERITAS, the current generation of space- and ground-based U.S.-led gamma-ray facilities;
1083 construction of the PINGU neutrino detector to lower the energy threshold to a few GeV and enable the
1084 determination of the neutrino mass hierarchy using atmospheric neutrinos; continued operation of the Auger
1085 and Telescope Array air shower arrays with upgrades to enhance the determination of the composition of

1086 the flux of cosmic rays around the GZK suppression region; construction and deployment of the JEM-EUSO
1087 mission aboard the International Space Station to extend observations of the cosmic ray flux and anisotropy
1088 well beyond the GZK region; and construction of a next-generation ultra-high-energy GZK neutrino detector,
1089 which will either detect GZK neutrinos (and constrain the neutrino-nucleon cross section at ultra-high energy)
1090 or exclude all but the most unfavorable parts of the allowed parameter space. A detailed census of present
1091 and proposed cosmic particle measurement facilities was produced during the Snowmass process.

1092 **Summary:** In synergy with the other frontier areas, the Cosmic Frontier provides to particle physics clear
1093 evidence for physics beyond the Standard Model; profound questions of popular interest; frequent new
1094 results, surprises, with broad impact; a large discovery space with unique probes; important cross-frontier
1095 topics; and a full range of project scales, providing flexible programmatic options. For each area of the
1096 Cosmic Frontier, the Snowmass study identified essential technologies and facilities, the advances required
1097 in theoretical models, and experiments with great promise. The largest projects are, appropriately and
1098 necessarily, international. The U.S. is still the leader in many areas of the Cosmic Frontier, but this field is
1099 evolving quickly and other regions with intensive interest in this physics are advancing rapidly.

1100 1.5 Theory

1101 This section summarizes the report of the Snowmass Theory Panel. The DPF constituted this panel with the
1102 goal of understanding both the scientific problems and opportunities of the next decade, and the challenges
1103 involved in sustaining a first-class theory program in the U.S.

1104 Theoretical physics has played a crucial role in particle physics since its earliest days. Theorists developed
1105 the basic framework in which we understand elementary particles: quantum field theory. This framework
1106 embodies Einstein's principles of special relativity and locality of interactions within the laws of quantum
1107 mechanics. It is extraordinarily successful. Theorists appreciated the role of symmetries as organizing
1108 principles for understanding data and clues to the nature of physical law. They developed calculational
1109 methods for quantum field theories, permitting the computation of scattering amplitudes, bound state
1110 masses, and numerous other quantities, often with extremely high precision. These developments combined
1111 to both produce and test the Standard Model.

1112 The discovery of a scalar particle at the LHC may well mark the completion of the Standard Model. This
1113 object is likely the Higgs boson of the simplest version of the theory. Theorists have played and will continue
1114 to play essential roles in firmly establishing the identity of this object. Its study at colliders requires not
1115 only great experimental ingenuity and persistence, but also an array of theoretical tools for calculating the
1116 rates for its production and decay. Just as crucial are techniques for the calculation of the large backgrounds
1117 arising from other Standard Model processes.

1118 While theoretical studies of quantum field theory and the SM have a long history, two areas have seen
1119 extensive progress in the last decade, and will continue to be the subjects of intense development. The first
1120 of these is perturbative methods for the calculation of scattering amplitudes. In recent years, calculations
1121 essential for collider physics, and previously believed essentially impossible, have been carried through using
1122 a range of new methods. These computations played a crucial role in the discovery of the Higgs boson, and
1123 are vitally important in searches for new physics. The development of new methods, and their application
1124 in LHC and other experiments, as well as in theoretical investigations, will remain a major activity in the
1125 field in the coming years.

1126 Another area of striking progress has been lattice gauge theory. This is the principal tool we have for the
1127 quantitative study of the strong interactions in processes at low energies. It is now possible to compute the

1128 spectrum of hadrons with high accuracy, and lattice computations have been crucial in the measurement of
1129 the properties of heavy quarks. Continuing improvements in calculational methods are anticipated in coming
1130 years. At Snowmass we heard about new applications of lattice methods, such as prospects for computations
1131 essential to any convincing interpretation of results from the upcoming muon $g - 2$ experiment.

1132 When we say the Standard Model may now be “complete” we mean that the theory is consistent to much
1133 higher energies. But there are strong reasons to believe that, at energies not much higher than those we probe
1134 today, there should be new phenomena. Theorists have been the drivers in formulating these questions, and
1135 proposing possible solutions. Among the questions are:

- 1136 1. Why are there vastly disparate mass scales in nature, such as the Planck mass and the weak scale?
- 1137 2. Why are neutrinos light?
- 1138 3. Are neutrinos their own antiparticles?
- 1139 4. What is the origin of the asymmetry between matter and antimatter?
- 1140 5. What is the identity of dark matter?
- 1141 6. What is the identity of dark energy?
- 1142 7. What is the origin of the curious pattern of quarks and leptons, and their masses?
- 1143 8. Do the forces unify?
- 1144 9. What modifications of our basic understanding are required to reconcile quantum mechanics and
1145 gravity?

1146 For each of these questions, theorists have proposed answers. Many of these are the subject of present or
1147 planned experimental searches.

1148 The Planck scale and the weak scale differ by at least 15 orders of magnitude. The difficulty of explaining
1149 the existence of such widely different scales is called the “hierarchy problem”. The large ratio of scales
1150 might be viewed as simply a “fact”, but within quantum field theory, this sort of hierarchy is generally quite
1151 unstable. (This is usually referred to as the “fine tuning” or “technical naturalness” problem.) Proposals
1152 to solve this problem all suggest physics at or near the TeV scale. Among the most explored of these
1153 is supersymmetry, a possible new symmetry of nature, which connects fermions and bosons. The LHC
1154 is actively searching for the new particles predicted by the supersymmetry hypothesis, and has excluded
1155 many popular models. Alternative proposals include the possibility that the Higgs particle is composite,
1156 or associated with phenomena in dimensions of space-time beyond the usual four. Over the next decade
1157 theorists will continue to explore these and other models, incorporating the constraints from experiment, or
1158 the results of discoveries.

1159 We know that the Universe consists largely of forms of energy not contained within the SM: dark matter
1160 and dark energy. The case for dark matter, a form of matter which behaves, essentially, like dust and
1161 interacts extremely weakly with ordinary matter, has become compelling in the past decade. Theorists
1162 have proposed several persuasive ideas for what the dark matter might be and how it was produced in the
1163 big bang. Supersymmetric models, in fact, naturally yield candidates (so-called weakly interacting massive
1164 particles, or WIMPs), which are automatically produced in roughly the right quantities. These particles are
1165 the subject of active experimental search at accelerators, deep underground, and in space. Theorists are
1166 working actively to survey the possible models and to understand exclusions and possible signals in ongoing
1167 and future experiments.

1168 But there are other candidates for the dark matter. Perhaps the most prominent of these is the axion.
1169 This particle was proposed to explain perhaps the largest remaining puzzle of the strong interactions: the
1170 conservation of CP. The equations of QCD include a parameter, θ , a pure number, which violates CP.
1171 Exquisite experiments set a limit on the size of any electric dipole moment of the neutron, and this, in turn,
1172 requires $\theta < 10^{-10}$. One possible explanation for this small number is a new particle, called the “axion”,
1173 whose dynamics adjust θ to a value close to zero. It turns out that if the axion exists, it is also a candidate
1174 for the dark matter. Theorists continue to refine the axion theory, exploiting developments in field theory
1175 and in string theory, and to explore its properties. The ADMX experiment at the University of Washington
1176 is currently searching for this particle and has a good chance to find it if it exists.

1177 Dark energy is equally mysterious. Representing about 70% of the energy budget of the Universe, this
1178 substance has *negative* pressure. Most theorists suspect that this is Einstein’s “cosmological constant”, and
1179 the data to date are consistent with this interpretation. But its value is very puzzling. Conventional ideas of
1180 quantum field theory suggest that there should be much more of it, and even more puzzling is the fact that
1181 its density is just such that it is becoming important in the current epoch of the Universe. These questions
1182 occupy the attention of many theorists, and there are a number of proposed answers, but it is safe to say that
1183 there is no compelling picture, and that this will certainly remain an active area of theoretical investigation
1184 for some time.

1185 Neutrinos are now known to have mass, and we know some features of their masses (mass matrix). Neutrinos
1186 are far lighter than other particles, and their masses are not accounted for within the SM itself. Theorists have
1187 identified two possible mechanisms to generate neutrino mass. One is associated with new particle. These
1188 particles might have enormous masses. In this case, neutrinos are their own antiparticles. Alternatively, there
1189 might be extremely light additional degrees of freedom. This is a question that can be tested experimentally,
1190 and which has several theoretical consequences. Understanding the neutrino masses and mixings will be a
1191 central part of both the experimental and theoretical particle physics programs over the next decade.

1192 Within the 5% of the energy budget which consists of ordinary protons and neutrons (baryons), there is a
1193 further puzzle: why is there matter at all, i.e., why didn’t the Universe emerge from the Big Bang with equal
1194 amounts of matter and antimatter? With the discovery of CP violation 50 years ago, it was recognized (first
1195 by Andrei Sakharov) that this is a question that can be addressed by science. Theorists have understood
1196 that the SM, however, does not violate CP sufficiently to account for the observed asymmetry; additional
1197 degrees of freedom (particles) are an essential component. They have put forward a number of proposals
1198 for how the asymmetry might arise. Some of the most compelling lie within the frameworks of theories of
1199 lepton mass (“leptogenesis”) and supersymmetry. These ideas might have observable consequences for the
1200 cosmos, such as the emission of gravitational waves, and for experiments at accelerators.

1201 Much of the history of particle physics has been tied up with the problem of “flavor”: the existence of three
1202 generations of quarks and leptons, and the features of their masses and mixings. Theorists have been central
1203 to this subject, proposing the idea of quarks and explaining the problem of mixing. In recent decades, they
1204 have developed theoretical tools to understand the behavior of heavy quarks, and a range of ideas for how
1205 the repetitive structure of quarks and leptons might emerge. These include ideas involving new symmetries,
1206 grand unification, string theory and extra dimensions. While many of these ideas are plausible, none are, as
1207 of yet, compelling in themselves, and these issues — dealing with the dynamics of heavy quarks and seeking
1208 an understanding of the basic issues of flavor — will be the focus of important activities in the next decade.

1209 The unification of forces is a long-standing dream. In the past few decades, theorists have put forward
1210 concrete ideas about how this might arise and proposals for experiments that could test the possibility.
1211 With supersymmetry, quite remarkably, the gauge couplings of the SM unify at a high energy scale, and
1212 many proposals have been put forward for an underlying explanation. A simple and compelling set of ideas
1213 of this type go by the name “Grand Unified Theories.” These elegantly enlarge the structure of the SM.

1214 The most dramatic consequence of all of these proposals is the prospect of proton decay, which has been
1215 the subject of extensive experimental search. Ideas of unification have also led to a rich set of theoretical
1216 questions, including the existence of magnetic monopoles.

1217 Beyond grand unified theories, the most ambitious attempt to unify the forces is associated with “superstring
1218 theory.” What is called string theory is part of a larger, only partly understood, structure which unifies
1219 Einstein’s general relativity and the other known forces in a quantum mechanical framework. While many
1220 questions are not yet answered, string theory has provided insight into longstanding questions in particle
1221 physics, including the unification of forces, the strong CP problem, dark matter and dark energy. It has also
1222 inspired much interesting phenomenology, such as that associated with large extra dimensions. This is an
1223 area that will continue to occupy a significant fraction of the community in the coming years, and which is
1224 likely to see significant additional progress.

1225 Much of the panel’s effort was devoted, as required by its charge, to examining structural issues in theoretical
1226 physics. The panel’s report deals at some length with questions of funding. The panel was concerned that
1227 the current budgetary climate at both NSF and DOE puts in jeopardy the research program that we have
1228 outlined above. Most prominent among our concerns was support for postdoctoral fellows and students.
1229 Students and postdocs are important drivers of research, and clearly represent the future of the field. The
1230 panel recommends keeping the current level of support for productive research groups, of roughly one postdoc
1231 and one graduate student per two PIs.

1232 The panel supports the program of comparative reviews, recently introduced by the DOE. This permits the
1233 agency to look critically at the support levels of individual theory groups, moving away from a model where
1234 funding levels were usually determined by making modest adjustments to historical levels of support. It
1235 permits the funding of new research groups and dropping groups which have become less productive (as the
1236 NSF has done historically). This is essential to adapting to the present funding climate.

1237 The panel understands the need to increase the fraction of the DOE budget devoted to projects, but argues
1238 that this has particularly severe consequences for theoretical physics. We proposed that the DOE consider a
1239 project category aimed at theory, and in particular designed to sustain a suitable population of postdoctoral
1240 fellows. One suggestion is the creation of “theory networks,” loosely modeled on networks established in
1241 Europe. The DOE would call for proposals to compete to establish such networks, with a lifetime of three
1242 to five years. The central topics of investigation would be determined by the institutions, but we envision
1243 that they might range from intense, phenomenological efforts in areas like neutrino physics, to investigations
1244 of more foundational issues in field theory and string theory.

1245 1.6 Accelerator Capabilities

1246 The Accelerator Capability study is a synthesis of individual workshops of six working groups plus the
1247 collective Snowmass meeting of all interested participants. Each group addressed major challenges foreseen
1248 for their respective class of accelerators in a pre-Snowmass meeting. The groups also considered a set of big
1249 questions regarding accelerator capabilities for the long-term future of high energy physics:

- 1250 1. How can one build a collider at the 10 – 30 TeV constituent mass scale?
- 1251 2. What is the furthest practical energy reach of accelerator-based particle physics?
- 1252 3. How would one generate ten or more megawatts of proton beam power?
- 1253 4. Can multi-megawatt targets survive? If so, for how long?

1254 5. Can plasma-based accelerators achieve energies and luminosities relevant to particle physics?

1255 6. Can accelerators be made an order of magnitude cheaper per GeV and/or per MW?

1256 The results of the workshops formed the basis for draft reports from the working groups that were discussed
1257 in the general Snowmass meeting to form this consensus summary.

1258 **Hadron colliders:** This working group focused on the evolution of the LHC and possible designs for a
1259 (much) higher energy proton collider (VLHC). The group considered: (1) how high a luminosity is possible
1260 for the LHC, (2) what are available increasing integrated luminosity without compromising experiments or
1261 detector survival, (3) how high an energy is possible in the LHC tunnel, (4) what impediments exist to
1262 designing a 100 TeV collider, and (5) what the associated accelerator research roadmap should be for hadron
1263 colliders.

1264 The priority recommendation of our study is full exploitation of the LHC. Doing so requires a strong LHC
1265 Accelerator Research Program sponsored by the Office of High Energy Physics that transitions to a US-
1266 LHC high luminosity construction project. During the project period we recommend continuing a focused,
1267 integrated, laboratory program that emphasizes the engineering readiness of technologies suitable for a 26
1268 TeV upgrade of the LHC or a machine of higher energy in a larger tunnel. The most critical technology
1269 development toward higher-energy hadron colliders is the next-generation high field Nb₃Sn magnets (limited
1270 to 15 Tesla) and adequate beam control technology to assure machine protection.

1271 The reach of an LHC energy upgrade is constrained by the limits of Nb₃Sn technology and by the absence of
1272 engineering materials with high-field properties beyond those of Nb₃Sn. Moreover, even doubling the LHC
1273 energy in the present tunnel introduces substantial issues of synchrotron radiation management. Radiation
1274 management will become very difficult as the synchrotron power on the beam tube reaches 5 W/m.

1275 To achieve energies beyond those of the LHC, the multi-laboratory study of VLHC remains valid. Snowmass
1276 has stimulated renewed effort on the VLHC in both the U.S. and Europe. American participation in the
1277 CERN-led international study for colliders in a large tunnel that will begin in 2014 will inform decisions to
1278 expand the reach of U.S. technology and guide research investments. The areas in which U.S. accelerator
1279 scientists can make the most valuable contributions are beam dynamics, superconducting magnets, vacuum
1280 systems, and machine protection.

1281 Long-term, innovative research will expand the technical options for any future hadron collider. Dipoles
1282 with operating fields beyond 15 T need new conductor elements such as small-filament, high-temperature
1283 superconductors in continuous kilometer lengths. Better conductors, innovative stress management, and
1284 novel structural materials will enable even higher-field magnets with greater temperature margin. With
1285 ever more stored energy in the beams, better understanding and modeling of beam dynamics is essential to
1286 control beam halos and lost beam particles. Machine protection and design of beam abort dumps for multi-
1287 GJ beams will be challenging. Other issues for research include effects of marginal synchrotron radiation
1288 damping, beam physics of the injection chain, effects of noise and ground motion, and options for interaction
1289 region design.

1290 **Energy-frontier lepton and photon colliders:** Our study welcomes the initiative for the International
1291 Linear Collider (ILC) in Japan. The ILC would begin as a 250 GeV Higgs factory with future expansion
1292 to 500 GeV. The U.S. accelerator community is capable of contributing to the ILC as part of a balanced
1293 U.S. particle physics program. As described in its Technical Design Report (TDR), the ILC is technically
1294 ready to proceed to construction. The TDR incorporates leadership U.S. contributions to machine physics
1295 and technology in superconducting RF (SRF), high-power targets for positron production, beam delivery,
1296 damping ring design, and beam dynamics such as electron cloud effects. Extending the ILC to 1 TeV would

1297 require lengthened linac tunnels and added cryomodules. It would use the original ILC sources, damping
1298 rings, beam delivery systems, and beam dumps.

1299 The excitement surrounding the Higgs boson discovery stimulated consideration of alternatives to a SRF
1300 Higgs factory. Concepts include a linear collider using copper linacs, a large circumference e^+e^- ring, a
1301 compact muon collider ring, a photon collider, and e^+e^- linear colliders based on wakefield acceleration
1302 techniques. These concepts span a broad range of technical readiness (from requiring demonstration of
1303 feasibility to having a detailed conceptual design) and timescales upon which a machine could be constructed.
1304 They also have varying energy reach from the 100s of GeV scale to the multi-TeV regime.

1305 It is natural to investigate whether a 250 GeV Higgs factory could fit in the LHC tunnel. This option is
1306 undesirable because it interferes with LHC operations and because the beam physics is highly constrained.
1307 Assessment of a circular collider in very large (of order 100 km) tunnel, with an energy reach up to about 400
1308 GeV, will be part of the CERN-led study of large colliders mentioned above. Such a machine is a substantial
1309 extrapolation from existing and past storage rings, albeit from a large experience base. Beamstrahlung at
1310 the interaction point strongly couples energy reach and luminosity. The luminosity would be largest at the
1311 Z peak, but fall rapidly as the center of mass energy increases. Should the ILC not go forward over the
1312 next decade and should the renewed interest in a very large circumference hadron collider be sustained, the
1313 possibility of a circular Higgs factory deserves extensive consideration.

1314 In a Higgs factory photon collider, two electron beams are accelerated to 80 GeV and converted to 63 GeV
1315 photon beams via inverse Compton scattering against low energy (3.5 eV), high intensity (5 J) laser pulses.
1316 The high-energy photon beams then collide to generate Higgs bosons through the s -channel resonance,
1317 $\gamma\gamma \rightarrow H$. A photon collider has the distinct advantage of requiring only an 80 GeV electron beam energy.
1318 Photon colliders could accompany proposed linear or circular colliders or be stand-alone facilities. The laser
1319 technologies overlap with those for laser wakefield accelerators.

1320 Muon accelerators could provide world-leading experimental capabilities at energies from the Higgs s -channel
1321 threshold at 126 GeV up to the multi-TeV scale. A circular muon collider, if feasible, could reach such
1322 energies, because the larger mass of the muon suppresses synchrotron radiation. As muons at rest have a
1323 lifetime of 1 μ s, they will decay in flight. The short muon lifetime demands that beam creation, manipulation
1324 and acceleration to high energy be done rapidly; high-gradient acceleration is essential. An Energy Frontier
1325 muon collider would necessarily be relatively compact. Even a 5 TeV collider would fit on the Fermilab site.
1326 Critical beam physics issues are: (1) cooling the muon phase volume by 10^6 , and (2) accumulating 10^{12}
1327 μ^+ and μ^- bunches in the collider. A vigorous, integrated R&D program toward demonstrating feasibility
1328 of a muon collider (Muon Accelerator Program) is highly desirable. The current funding level is, however,
1329 insufficient for timely progress. Development of a muon collider capability would be closely connected with
1330 Intensity Frontier accelerators such as intense neutrino sources.

1331 The Compact Linear Collider (CLIC) concept is based on 100 MeV/m copper linac technology; it would
1332 stretch 50 km for a 3 TeV collider. CLIC would be powered by two high-current drive beams running parallel
1333 to the colliding beams through a sequence of power extraction and transfer structures, where they produce
1334 short, high-power RF pulses that are transferred into the accelerating structures. The practical energy reach
1335 depends on control of wakefields and on the accelerating gradient in industrialized accelerator sections. U.S.
1336 national laboratories have substantial expertise in CLIC technologies.

1337 Yet another approach for multi-TeV energies proposes to use wakefields in plasmas or dielectric structures
1338 driven either by beams or lasers to achieve accelerating fields of 10 to 100 GeV/m. Many feasibility and
1339 practicality issues remain: positron acceleration, multi-stage acceleration, control of beam quality, and
1340 plasma instabilities at tens of kHz repetition rate. All variants require an integrated proof-of-principle
1341 test. The U.S. is a world leader in these strong physics programs at the frontier of accelerator science.

1342 **High-intensity proton sources for neutrinos, muons, and rare processes:** Requirements for Intensity
1343 Frontier experiments are more diverse than for the energy frontier. Therefore, this study addressed a set of
1344 structured questions: (1) What secondary beams are needed for Intensity Frontier experiments? (2) What
1345 proton beams could generate such beams? (3) Can these proton beams be made by existing machines? (4)
1346 What new capabilities are needed? (5) What accelerator and target research is needed to realize the new
1347 capabilities? The study surveyed particle physics requirements for secondary beams, including beams of
1348 neutrinos, kaons, muons, and neutrons. Experiment advocates supplied nineteen secondary beam requests.
1349 From these the study group derived primary proton beam characteristics.

1350 The common characteristics required are average beam power, with more than 1 MW delivered, and a
1351 flexible, experiment-dependent time structure. Beam requirements were compared with 20 existing proton
1352 beam-lines and 14 planned upgrades. The overarching conclusion is that the next generation of intensity
1353 frontier experiments requires beam intensities and timing structures beyond the capabilities of any existing
1354 accelerator.

1355 Fermilab's proposed, multi-stage, Project X would yield a world-leading facility based on a modern multi-MW
1356 superconducting proton linac capable of injecting into the Fermilab Main Injector. The linac would deliver
1357 a flexible on-demand beam structure that could serve multiple experiments over an energy range 0.25–120
1358 GeV. The linac would provide a platform for future muon facilities including nuSTORM, a neutrino factory,
1359 and a muon collider. A complete, integrated Reference Design Report identifies technical risks that will be
1360 mitigated in a structured research program already underway.

1361 The DAE δ ALUS collaboration proposes multiple sources of decay-at-rest anti-neutrinos for short-baseline
1362 oscillation experiments. This project has narrower experimental scope than Project X. DAE δ ALUS would
1363 use three multi-MW H₂⁺ cyclotrons and target stations located about 2 to 20 km from a large hydrogenous
1364 detector. The experiment would measure CP violation in a way that is complementary to the LBNE
1365 experiment. The first stage of DAE δ ALUS would be IsoDAR, a compact 60 MeV cyclotron located only
1366 15 m from the KamLAND detector, that would make a definitive search for one or two sterile neutrinos.
1367 This international collaboration has strong connections with commercial cyclotron industries.

1368 Another exciting possibility is nuSTORM, Neutrinos from STORed Muons. This would be a first step toward
1369 a long-baseline neutrino factory capability. The nuSTORM muon storage ring would send well-characterized
1370 neutrino beams to detectors at 50 m and 1900 m for a sterile neutrino search and neutrino cross-section
1371 measurements.

1372 A common research issue for Intensity Frontier capabilities is the injection system, composed of low-
1373 emittance, high-current ion sources with effective beam choppers. Control of space-charge forces is important
1374 for preserving beam quality. Understanding and limiting beam loss is a dominant operational issue requiring
1375 adequate simulation of halo formation, efficient beam collimation, and very high-efficiency extraction.

1376 High-power targets are a difficult challenge that limits facility performance. The principal underlying damage
1377 mechanisms of the target materials are atom displacements and gas production. Particulars depend on
1378 primary beam characteristics, target material, operating temperature, and the duty factor of the accelerator.
1379 Unfortunately, one cannot directly translate experience with nuclear reactors to estimate the performance
1380 of targets with high-energy beams from experience with nuclear reactors. Details of target behavior and
1381 failure mechanisms are a mesoscale problem that is difficult to simulate. Computed radiation effects in
1382 inhomogeneous materials subject to time-varying irradiation need validation with controlled, instrumented
1383 in-beam tests.

1384 **High-intensity electron and photon beams:** This working group addressed two major questions: (1)
1385 What capabilities at heavy flavor factories are required to realize the full range of physics opportunities? (2)

1386 What are new physics opportunities using high power electron and positron beams? The relevant technologies
1387 exploit strong synergy with light sources and damping rings.

1388 SuperKEKB, a super-high-luminosity B-factory, is an upgrade to the KEKB B-factory currently under
1389 construction in Japan with commissioning to commence in January 2015. To achieve the target luminosity
1390 of $8 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$, a forty-fold increase over that of KEKB, the SuperKEKB beam currents will be
1391 approximately twice as high as used at KEKB, and vertical bunch sizes at the collision point about 20 times
1392 smaller than those achieved at KEKB. Greater U.S. collaboration would strengthen the SuperKEKB project
1393 and might enable even higher luminosity well in excess of $10^{36} \text{ cm}^{-2}\text{s}^{-1}$.

1394 Two super tau-charm factories have been proposed: one at Frascati (Tor Vergata) in Italy and one at Novosi-
1395 birsk in Russia. Both machines are two-ring, symmetric-energy machines, with provisions for longitudinally
1396 polarized beams.

1397 The DarkLight experiment will use the high-intensity electron beam of the JLab FEL, impinging on a
1398 hydrogen target, to search for gauge bosons associated with “dark force” theories. It might also be possible to
1399 use an intense, low-emittance positron beam, impinging on a plasma target, to generate sufficient muon/anti-
1400 muon pairs to provide a source beam for a future muon collider without the need for a separate muon cooling
1401 stage.

1402 **Electron-ion colliders:** Several future electron-ion colliders have been studied in recent years. All would
1403 be based at an existing accelerator facility. The collider configurations include both ring-ring and linac-ring
1404 options. Center of mass energies range from 14 GeV to 2000 GeV. Most of the collider concepts share several
1405 enabling technologies. SRF cavities must be able to operate with high average and high peak beam currents,
1406 providing effective damping of high-order modes. The cryomodule design must be consistent with containing
1407 high beam power. Hadron beam transverse and longitudinal emittances must be small to achieve high
1408 collider luminosity. Therefore, the designs with medium hadron energy call for the application of powerful
1409 cooling techniques.

1410 The low β^* interaction region designs for all proposed colliders require strong focusing of beams at the
1411 collision point and fast separation of beams after the collision. The synchrotron radiation fan produced by
1412 electrons in the focusing magnets must be kept from hitting the the beam pipe in the vicinity of the detectors
1413 and inside superconducting magnets.

1414 The linac-ring designs utilize a polarized electron source, with an average current ranging from 6 mA to
1415 50 mA. The linac-ring scheme introduces non-standard beam-beam effects, which must be explored to
1416 understand the limits on the luminosity and the beam parameters. Other shared technologies include
1417 techniques to preserve beam polarization. Spin matching and the harmonic correction techniques have
1418 to be investigated for ring-ring colliders to minimize the beam depolarization due to synchrotron radiation,
1419 especially in the presence of spin rotators and solenoidal detector magnets.

1420 **Accelerator technology test beds:** We identified a broad range of existing and needed test capabilities
1421 for proposed frontier accelerators. The first category of test facilities permits testing beam physics or
1422 accelerator components essential to manage technical risks in planned projects. A second category would
1423 integrate accelerator systems to provide proof-of-practicality tests. The third category provides tests of
1424 physics feasibility of concepts and/or components. This study identified 35 existing facilities in the U.S.
1425 and overseas, both with and without beam testing capability. Although these facilities provide substantial
1426 readiness to move forward with the highest priority accelerators for particle physics, the long range future
1427 of particle physics needs a few additional dedicated test capabilities in the near term.

1.7 Underground Laboratory Capabilities

Some of the most compelling experiments in particle physics can only be done at underground facilities. Searches for dark matter and neutrinoless double beta decay and neutrino experiments using solar, reactor, atmospheric, and supernova neutrinos and neutrinos from accelerators all require underground facilities and capabilities.

Underground facilities are located in North America, Europe, Asia, and Antarctica (in ice). New underground facilities have become operational in all of these regions in the last few years. The world-wide particle physics community plans to expand underground capabilities over the next years, primarily outside the United States. If all of these plans are realized, general-purpose space for underground experiments will roughly double by the end of the decade. The expansion would include major new facilities to host reactor experiments at moderate depths, and a new class of very large facilities for long-baseline and atmospheric neutrino experiments, proton decay, and other physics.

Plans for expansion or continuation of underground facilities in the United States are less developed. Currently, there are no plans with approved federal funding for expansion of underground capabilities at the four underground sites located in the United States. The Long-Baseline Neutrino Experiment (LBNE) has provisional approval to be located on the surface at the Sanford Underground Research Facility (SURF) in South Dakota, but design work is underway in anticipation of achieving a global collaboration to allow LBNE to be sited deep underground at SURF. The LBNE physics community expressed strong support for the deeper site during the Snowmass process.

All of the next generation (G2) dark matter experiments can be accommodated by existing or planned underground facilities, assuming no reduction in these facilities for the rest of the decade. Only one of these experiments is planned to be located at a U.S. facility. Several neutrinoless double-beta-decay experiments are already under construction at existing underground facilities, one of which is in the U.S. next-generation (ton scale) neutrinoless double beta decay experiments can likely be accommodated by existing and planned facilities, but will face competition for underground laboratory space from dark matter experiments.

Detectors for reactor experiments with baselines greater than 100 m require medium-depth underground laboratories. Future reactor experiments are being planned overseas based on funding commitments from the host countries.

The flagship of the international neutrino effort is the search for CP violation in the lepton sector, which requires a massive detector and a very intense neutrino beam. There are other motivations for constructing this massive detector underground. The search for nucleon decay is one of the most important topics in particle physics. Atmospheric neutrinos, observable in a large underground detector, may be sensitive to all of the currently poorly known neutrino oscillation parameters. The spectacular neutrino burst from a nearby supernova event would be detected at no additional cost if the detector is underground, but such detection is very difficult on the surface. Some of the same detectors that would be used for long-baseline neutrino experiments could be used to advance the search for CP violation, nucleon decay, the study of atmospheric neutrinos and other physics if the detector is located underground. This is the plan for Hyper-K (Japan) and LBNO (Finland). It would be a lost opportunity if this condition cannot be satisfied with LBNE.

Experimental needs for materials assay and storage outstrip the capability of existing facilities, and space for such work should be reserved at new facilities. In addition, underground space should be reserved for small prototype testing and generic R&D. There is enough space at U.S. facilities to meet future needs if the existing underground labs are maintained.

1470 As the scale and cost of underground experiments grows it will become even more important to maintain
1471 open competitive access to underground laboratories. The best way for the governments to support the
1472 international system of underground experiments is for each major country (or region) to support at least
1473 one major underground laboratory capable of hosting forefront experiments. It is not clear whether it would
1474 be possible to sustain this international support if one country chose to take a major role in the research
1475 without supporting any facility.

1476 Our conclusions are:

- 1477 1. We should locate LBNE underground to realize its full science potential. This step would also provide
1478 a natural base for additional domestic underground capabilities at SURF in the future.
- 1479 2. The U.S. has leading roles in many of the future dark matter, neutrinoless double-beta-decay and
1480 neutrino experiments.
- 1481 3. More coordination and planning of underground facilities (overseas and domestic) is required to
1482 maintain this leading role, including use of U.S. infrastructure.
- 1483 4. Maintaining an underground facility that can be expanded to house the largest dark matter and
1484 neutrinoless double-beta-decay experiments would guarantee the ability of the U.S. to continue its
1485 strong role in the worldwide program of underground physics.

1486 1.8 Instrumentation

1487 The search for answers to fundamental questions in the field of particle physics has always been intimately
1488 tied to the development of innovative technologies or significant advancements in existing technologies.
1489 The particle physics community has a long history of inventing detectors based on new technologies to
1490 address the science needs and advancing these technologies to large-scale reliable use. For many decades the
1491 instrumentation needs of particle physics have motivated university faculty, national laboratory staff, and
1492 industrial scientists to develop the technological foundations and invent the detectors responsible for many of
1493 the important particle physics discoveries. The field of particle physics is generally regarded as an incubator
1494 of innovation in instrumentation. Moreover, driven by the needs of particle physics, the technology developed
1495 for accelerators and detectors has historically benefited many other fields of the physical and applied sciences
1496 and medicine.

1497 Modern particle physics experiments and high-energy accelerators put increasingly extraordinary demands
1498 on sensors, sensor readout electronics, precision engineering, and data acquisition and management, often
1499 incorporated into detector systems of very large scale. In addition, the scientific approaches are broadening to
1500 include lower-energy, high-intensity and ultra-low-background experiments, including experiments deploying
1501 very large volume detectors to study very rare processes, and also experiments that study fundamental
1502 properties of the cosmic energy and matter with increasingly greater precision.

1503 Focusing only on scaling up existing technologies to larger experiments or carrying out detector R&D only
1504 when it is needed is tempting in tight budgetary times, but this is counter to the successful approach that
1505 has been followed up to now. Instead, more fundamental innovation and development of new approaches
1506 is necessary to make experiments feasible or economically viable. An appropriate investment in a detector
1507 R&D program will be required to enable the science goals and meet the budgetary challenges. This program
1508 would develop, over intermediate and long time frames, new tools and technologies that are both cost-
1509 effective and have an enhanced physics reach. It will allow the field to continue to carry out flagship domestic
1510 experimental research and have leadership roles in off-shore experimental projects, while the development

1511 of new, transformative detection capabilities will ensure an affordable and healthy experimental particle
1512 physics research program in the future. The major challenge for instrumentation is to structure the current
1513 advanced detector R&D program such that it will enable the United States to continue to maintain scientific
1514 leadership in many key areas of a broad international experimental program in particle physics.

1515 The instrumentation needs of planned and proposed future experiments across the field were surveyed in
1516 two joint preparatory meetings, in several dedicated topical workshops and in joint sessions with the Energy,
1517 Cosmic, and Intensity Frontiers during the Snowmass study. In addition, nearly one hundred white papers on
1518 instrumentation were submitted, covering a broad range of topics. In all cases instrumentation development
1519 was considered central to progress. An overview of the current and planned programs of experiments at the
1520 various physics frontiers and specific detector needs is provided in full Instrumentation report. This survey
1521 of the whole experimental program of particle physics identifies key issues in instrumentation within the next
1522 decade and beyond. It also gives a picture of the opportunities that exist to establish an instrumentation
1523 program that can satisfy both the needs of particle physics and at the same time create an environment
1524 of innovation to benefit other fields of science. This summary articulates a vision for instrumentation that
1525 will enable execution of a broad, targeted experimental program within the fiscal realities of our time and
1526 identify areas where the U.S. can take a leadership position.

1527 Ideally, detector development should take place at various levels of risk and various maturation time scales.
1528 Detector R&D carried out within existing experiments is, by necessity, project driven and provides a low-risk
1529 path to relatively incremental improvements to existing technologies. Detector R&D that is motivated by
1530 common needs among various experiments is generally longer-term, can be higher-risk, and adds value to
1531 multiple experimental areas at the same time. This type of R&D can lead to incremental or significant
1532 improvements in cost reduction, scientific reach, or both. At the highest level of impact and risk is long-term
1533 detector R&D leading to transformative changes in cost reduction, increase in scientific reach or both, across
1534 a significant part of the experimental program. This is the kind of high-risk, high-reward detector R&D that
1535 has the potential to lead to scientific breakthroughs. Underpinning these R&D efforts is the urgent need for
1536 training the next generation of instrumentation experts, without which there can be no long-term future.

1537 Major technological advances based on a better understanding of the underlying science are also occurring in
1538 other scientific disciplines such as materials science, photonics and nanotechnology. Many of these advances
1539 have the potential to lead to transformational new technologies for particle physics detectors. Particle physics
1540 should continue to exploit and pursue the technology advances in other experimental scientific disciplines,
1541 which could contribute to opportunities for innovation and the development of transformative technologies
1542 for particle physics.

1543 A healthy national instrumentation program must provide a balance between evolutionary and revolutionary
1544 detector development while training the next generation of experts. We have five recommendations for critical
1545 elements of this national program:

- 1546 1. Support detector R&D with clearly identified areas of detector development based on the strengths of
1547 the community.
- 1548 2. Achieve an appropriate balance between evolutionary and revolutionary detector R&D, i.e., an appropriate
1549 “portfolio of risk,” and build expertise in new and innovative technologies that can be applied
1550 to the design and construction of novel, cost-effective particle physics detectors.
- 1551 3. Develop a process for optimizing the use of existing university, national laboratory, and industrial
1552 resources, to grown and retain local technical expertise at universities and laboratories, and to identify
1553 incentives and mechanisms for improving detector R&D collaborations and equipment sharing among
1554 universities, national laboratories and industry.

- 1555 4. Create opportunities for attracting, and providing careers for, particle physicists with interest in,
1556 and outstanding capability for, innovative detector design and development. This community of
1557 experimental physicists will preserve the background and skills needed to design and build future
1558 generations of particle physics experiments.
- 1559 5. Provide mechanisms for identifying and transferring appropriate technologies developed in other sci-
1560 entific disciplines to particle physics and for transferring applicable technologies developed in particle
1561 physics to other science disciplines, such as nuclear physics, basic energy sciences, and related branches
1562 of science, medicine, and national security.

1563 The Snowmass process provided broad input and guidance from the particle physics community in identifying
1564 these crucial elements of a national technology and instrumentation program. A previous DPF Task
1565 Force on Instrumentation led to the creation of CPAD, the Coordinating Panel for Advanced Detectors,
1566 which is intended to act as the advocate of this program, promote its merits, and provide venues for
1567 regular presentation of results. It also brings different groups of technical experts together to make the
1568 community aware of developments elsewhere. CPAD can also coordinate between the funding agencies and
1569 the instrumentation developers by providing information about instrumentation needs and ongoing activities.

1570 A crucial enabling element for an instrumentation program is the development of expert physicist manpower.
1571 An investment in younger physicists who work on instrumentation is critical to support a long-term particle
1572 physics program. It is noted that only under exceptional circumstances are doctorate degrees awarded
1573 by U.S. physics department for a Ph.D. thesis based on instrumentation, whereas this is commonplace in
1574 Europe. Furthermore, early specialization in instrumentation is strongly disadvantaged given the emphasis
1575 on physics analysis to obtain faculty positions. This has to change for the field to remain viable.

1576 Other essential enablers are national laboratory resources, unique facilities such as test beams, and targeted
1577 funding aimed at a specific problem. The U.S. national laboratories are a resource of unique importance.
1578 Their breadth, because of their multidisciplinary scientific nature, provides cross-fertilization of useful
1579 technologies from nuclear physics, basic energy science, materials research, engineering, chemistry, and
1580 computer science. Many U.S. national laboratories have close links to top-ranked universities, which are also
1581 centers for multi-disciplinary innovation and ideas. University groups have greater difficulty maintaining
1582 long-term technical and engineering resources, while the national laboratories naturally maintain these as
1583 consequence of their missions. By combining the intellectual and manpower capabilities of universities with
1584 the resources of the national laboratories and the product development capabilities of industry, the U.S. can
1585 continue to confront and overcome many of the technological challenges of future particle physics experiments.

1586 The United States has several complementary test beam and irradiation facilities. These facilities are a
1587 critical component of instrumentation work, and there is a separate Snowmass report about them. The
1588 Snowmass process did not allow for a proper evaluation of all facilities for instrumentation, and it is
1589 recommended that CPAD finalize this process. Adequate support for these facilities is essential for a healthy
1590 detector R&D program.

1591 Another enabler to encourage transformative innovation in instrumentation would be to initiate a new
1592 program, outside the existing funding, for long-term investment in more speculative but potentially high
1593 impact research motivated by a set of “grand challenges.” Once identified, such grand challenges would
1594 be effective in focusing the creative power of the instrumentation community on problems that have the
1595 potential for large payoff. Areas where existing technologies would be cost-prohibitive for meeting the
1596 goals of future experiments are good candidates for new initiatives. These grand challenges should be
1597 issued nationally, and cross-disciplinary collaboration should be strongly encouraged. This would allow the
1598 program to take advantage of the tremendous progress and breakthrough advances that have been made
1599 in areas of science outside the field of particle physics that could prove very valuable for the development

1600 of future instrumentation. Because of the cross-disciplinary aspect, close collaboration between universities
1601 and national laboratories will be a key component. Funding would be subject to proposal review, but should
1602 be at a substantial level for a period of at least three years. CPAD could be engaged to identify the set of
1603 grand challenges that would define the program.

1604 During the Snowmass Instrumentation discussions, a number of instrumentation areas were recognized as of
1605 strategic importance. These areas all focus on major technological barriers that stand in the way of reaching
1606 the science goals. Some of them have the potential to deliver very cost-effective instrumentation methods
1607 and provide breakthrough new technology. The choice of these areas was guided by their physics impact and
1608 existing strengths and capabilities in the country. Consideration was given to the technology's usefulness to
1609 other branches of science. Although some of these instrumentation themes seem very challenging, it is likely
1610 that many, if not all, can be realized with a dedicated instrumentation effort. The main strategic areas that
1611 have been identified are described below.

1612 **ASICs:** The use of Application Specific Integrated Circuit (ASIC) electronics and interconnect development
1613 is often critical to enable an experiment. A number of factors make ASICs essential to particle physics, such
1614 as small physical size, high channel density, ability to integrate a variety of function blocks, low power
1615 dissipation and radiation tolerance. Examples of areas where ASIC-related R&D is required are high-speed
1616 waveform sampling, pico-second timing, high-rate radiation tolerant data transmission, low temperature
1617 operation, low power and 2.5D and 3D assemblies. The field of particle physics has spearheaded the use
1618 of ASICs, but there is a growing need and adoption by other disciplines. ASIC development provides
1619 an excellent opportunity to work more closely with other branches of science by trying to address their
1620 instrumentation needs. A report summarizing a workshop held earlier in 2013 to look at the ASIC needs for
1621 HEP was submitted to the Snowmass proceedings.

1622 **Calorimetry:** The measurement of the total energy of electrons and jets lies at the core of experiments
1623 at the Energy and Intensity Frontiers. Projects designed to search for very rare processes or to make
1624 precision measurements are in need of more precise, faster, and more cost-effective methods to perform these
1625 calorimetric measurements.

1626 **High speed data acquisition:** Experiments are required to handle huge interaction rates to acquire,
1627 transport, process and retain the events of interest, preserve the accuracy of the measurements of intrinsic
1628 particle properties, and uncover signatures of new physics. More intelligent trigger and data acquisition
1629 systems are needed to enable higher statistics experiments.

1630 **Large volume detectors:** The study of neutrino properties and their interactions and the search for dark
1631 matter require large volume detectors at underground facilities. Innovative technologies that allow scaling in
1632 a cost-effective way with increased sensitivity are required to enable the spectroscopy of these fundamental
1633 particles. A coherent research program in low-radioactive materials and assay is required.

1634 **Photodetectors:** A multitude of physics processes can be studied by measuring photons with wavelengths
1635 ranging from mm to nm. Instruments used to study these photons are based on a range of materials ranging
1636 from superconductors to semiconductors, from alkali metals to crystals. The development of large arrays
1637 with improved spectral sensitivity, energy and time resolution, and excellent background rejection would
1638 truly revolutionize future experiments.

1639 **Pixelated sensors:** High granularity has become a requirement for many of our detectors. Often the
1640 higher density results in performance compromises. The development of new technologies designed to deal
1641 with the higher density, while avoiding these compromises and improving overall performance, is essential
1642 for future experiments. These include sensors with a greater degree of pixelation, radiation hardness, high
1643 speed, and built-in intelligence to carry out a number of operations, including hit time-stamping, clustering
1644 and recognizing hit correlations, that can affordably be deployed in large areas.

1645 **Power and mass:** Especially at the Energy Frontier, experiments are characterized by high radiation, huge
1646 interaction rates and serious constraints on power and mass budget. Better low-mass structural materials
1647 that are strong and stable, including materials with ultra-low intrinsic radioactivity, would benefit a broad
1648 spectrum of future experiments. The design of electrical power distribution and cooling systems seems
1649 mundane, but it can severely limit the physics reach of current experiments. These systems must deliver
1650 services with low mass in a high radiation and magnetic field environment. Innovative solutions are critically
1651 important for next generation experiments.

1652 The particle physics technology and instrumentation program described here requires a multi-year commit-
1653 ment from the funding agencies. The funding required to meet short-term financial obligations to sustain an
1654 existing particle physics research program puts enormous pressure on funds earmarked for long-term, generic
1655 detector development. In spite of these pressures, a stably and adequately funded generic instrumentation
1656 program will ensure that the field invests in its future and establishes a foundation for a competitive, healthy
1657 program in the long term.

1658 1.9 Computing

1659 Computing has become a major component of all particle physics experiments and in many areas of
1660 theoretical particle physics. The Computing study group established subgroups covering user needs and
1661 infrastructure. The study considered user needs for experiments at the Energy and Intensity Frontiers,
1662 and the combined needs of Cosmic Frontier experiments, astrophysics and cosmology. Theory subgroups
1663 covered accelerator science, astrophysics and cosmology, lattice field theory, and perturbative QCD. Four
1664 infrastructure groups examined trends in computing to predict how technology will evolve and how it will
1665 affect future costs and capabilities. These groups focused on distributed computing and facility infrastruc-
1666 tures, networking, data management and storage. and software development, personnel, and training. They
1667 identified critical technology needs for particle physics that might require the DOE or NSF to fund research
1668 in computer science and technology.

1669 During the period between the Community Planning Meeting at Fermilab and the Community Summer
1670 Study meeting at the University of Minnesota, the Computing groups were actively engaged with the other
1671 frontiers to learn of their plans and estimate their computing needs. The infrastructure groups engaged
1672 with vendors, computer users, providers and technical experts to predict trends in computing, networking,
1673 storage, and software development, including considerations of costs, capacities and speeds. Two days of
1674 parallel sessions at the Minnesota meeting were devoted to discussions across the the subgroups, to finalize
1675 subgroup findings, and to identify common trends and needs.

1676 Progress in particle physics experiment and theory will require significantly more computing, software
1677 development, storage and networking. Different projects stretch future capabilities in different ways, but
1678 there are many common needs among the different areas of particle physics. In the future more commonality
1679 and community planning would aid in moving ahead in the most efficient manner. This requires careful and
1680 continuing review of the topics we studied, in particular, user needs and capabilities of current and future
1681 technology. For many years, the particle physics community has been a great source of computing innovation
1682 and expertise. It is essential to leverage those assets through wider sharing of knowledge throughout the
1683 experimental and theoretical communities. We should be open to bi-directional sharing of expertise with the
1684 entire scientific community.

1685 The experimental program relies for the most part on distributed high-throughput computing (HTC).
1686 Simulation, data analysis, and reconstruction of individual events are independent of each other, so that
1687 groups of events can be assigned to hardware in different locations. Results are combined when all event

1688 groups are done. This distributed computing model was pioneered by the Energy Frontier experiments. It
1689 relies on a distributed infrastructure of computing centers as part of the Open Science Grid in the U.S. and
1690 extending across the globe. Theoretical computing and simulation needs are more commonly addressed by
1691 high-performance computing (HPC), in which thousands to hundreds of thousands of tightly coupled CPUs
1692 are working simultaneously on a single problem. These resources are provided mostly through DOE and
1693 NSF supercomputing centers.

1694 One issue for those applications that traditionally rely on HTC for their data-intensive computing is to what
1695 degree they can or should use national supercomputer centers, which have traditionally been designed for
1696 HPC usage. Work is proceeding to make these HTC applications run on HPC, and to interface HPC centers to
1697 the HTC workload and data management infrastructures. Also, traditional HPC applications are developing
1698 in the direction of more data-intensive science, which, however, is currently not a good match to existing and
1699 next-generation HPC architectures. Computational resources will have to address the demands for greatly
1700 increasing data rates, and the increased needs for data intensive computing tasks like data analytics, for
1701 comparing large samples of simulations and observational data.

1702 Another pressing issue facing both HTC and HPC communities is that processor speeds are no longer
1703 increasing exponentially, as they were for at least two decades. Instead, new chip architectures provide
1704 multiple cores. Thus, we cannot rely on new hardware to run serial codes faster, and we must parallelize
1705 codes to increase application performance. In addition to multi-core chips, there are accelerators such as
1706 graphical processing units (GPUs) and many-core chips such as the Intel Xeon Phi. Computing resource
1707 needs for Energy Frontier experiments used to scale roughly with the rate that processor speeds increased,
1708 following Moore's law. Future advances will require full use of multiple-core and many-thread architectures.
1709 Scaling of disk capacity and throughput is of significant concern in the storage area, since per-unit capacities
1710 are no longer increasing as rapidly.

1711 These changes in chip technology and high performance system architectures require us to develop parallel
1712 algorithms and codes, and to train personnel to develop, support, and maintain them. Different subgroups
1713 are at different stages in their efforts to port to these new technologies. Lattice QCD, for example, started its
1714 GPU porting efforts in 2008 and has had code in production for some time, particularly for matrix inversions
1715 of the lattice Dirac operator; however, there are other parts of the code that are still only running on CPUs.
1716 Cosmological simulations have exploited GPUs since 2009 and some codes have fully incorporated GPUs in
1717 their production versions, running at full scale on hybrid supercomputers. Accelerator science is also actively
1718 porting codes to GPUs. Some of the solvers and particle-in-cell infrastructures have been ported and very
1719 significant speed-ups have been obtained. The perturbative QCD community has also started using GPUs.

1720 These trends lead to vastly increasing code and system complexities. For example, only a limited number
1721 of people in the field can program GPUs. In this and other highly technical areas, developing and keeping
1722 expertise in new software technologies is a challenge, because well-trained personnel and key developers are
1723 leaving to take attractive positions in industry. Continued training is an important aspect. Training materials
1724 are now provided by some of the national supercomputing centers, and by summer schools organized by,
1725 among others, the Virtual School of Computational Science and Engineering. We must examine whether
1726 these provide the right training for our field and whether the delivery mechanisms are timely. On-line media,
1727 workbooks and wikis are suggested to enhance training. Another area of common concern is the career path
1728 of those who become experts in software development and computing. It is useful to help young scientists
1729 learn computing and software skills that are marketable for non-academic jobs, but it is also important that
1730 there be career paths within particle physics, including tenure-track jobs, for those working at the forefront
1731 of computation.

1732 **Energy Frontier** experiments already experience computing limitations that limit the amount of physics
1733 data that can be analyzed. The planned upgrades to the LHC energy and luminosity are expected to result

1734 in a ten-fold increase in the number of events and a ten-fold increase in event complexity. Effort has begun to
1735 increase code efficiency and parallelism in reconstruction software and to explore the potential of GPUs. The
1736 experiments are considering saving more raw events to tape and only reconstructing them selectively. The
1737 LHC produces about 15 PB of raw data per year now, but by 2021 the rate may rise to 130 PB. Attention
1738 needs to be paid to data management and wide-area networking, to assure that network connectivity does not
1739 become a bottleneck for distributed event analysis. It is important to monitor storage cost and throughputs.
1740 More than half of the computing cost is now for storage, and in the future it may become cost-effective to
1741 recompute certain derived quantities rather than storing them.

1742 **Intensity Frontier** experiments have combined computing requirements on the scale of a single Energy
1743 Frontier experiment, but they form a more diverse set than those of the Energy Frontier. Our survey found
1744 that there is significant overlap in different experiments' needs. Sharing of resources across experiments, as in
1745 the Open Science Grid, is a first step in addressing peak computing resource needs. Continued coordination
1746 of software development between these experiments will allow for efficiently developed coding infrastructure.
1747 Leveraging the data handling experience and expertise of the Energy Frontier experiments for the diverse
1748 Intensity Frontier experiments would significantly improve their ability to reconstruct and analyze data.

1749 **Cosmic Frontier** experiments will greatly expand their storage needs with the start of new surveys and
1750 the development of new instruments. Current data sets are about 1 PB, and the total data set is expected
1751 to be about 50 PB in ten years. Beyond that, in 10–20 years data will be collected at the rate of 400 PB/yr.
1752 On the astrophysics and cosmology theory side, some of the most challenging simulations are being run on
1753 supercomputers. Current allocations for this effort are approximately 200M core-hours annually. Very large
1754 simulations will require increasing computing power. Comparing simulations with observations will play a
1755 crucial role in interpreting experiments, and simulations are also needed to help design new instruments.
1756 There are very significant challenges in dealing with new computer architectures and very large data sets, as
1757 described above. Growing archival storage, visualizing simulations, and allowing public access to data are
1758 also issues that need attention.

1759 **Accelerator Science** is called on to simulate new accelerator designs and to provide near-real-time simula-
1760 tion feedback for accelerator operation. Research into new algorithms and designs has the potential to bring
1761 new ideas and capabilities to the field. It will be necessary to include additional physics in codes and to
1762 improve algorithms to achieve these goals. Production runs can use from 10K to 100K cores. Considerable
1763 effort is being expended to port to new architectures, especially to address the real-time requirements.

1764 **Lattice Field Theory** calculations rely on national supercomputer centers and hardware purchased for
1765 the USQCD Computing Project. Allocations at supercomputer centers have exceeded 500 M core-hrs this
1766 year, and resource requests will go up by a factor of 50 by the end of this decade. This program provides
1767 essential input for interpretation of a number of experiments, and increased precision will be required in the
1768 future. For example, the b quark mass and the strong coupling α_s will need to be known at the 0.25% level,
1769 a factor of 2 better than now, to compare upcoming ILC Higgs observations with SM predictions. Advances
1770 in the calculation of hadronic contributions to muon $g - 2$ will be needed for interpretation of the planned
1771 experimental measurement.

1772 **Perturbative QCD** is essential for theoretical understanding of collider physics rates. Experts in perturba-
1773 tive QCD computation ported codes to the HPC centers at NERSC and OLCF, and to the Open Science
1774 Grid. They have also been benchmarking GPU codes and finding impressive speed-up over a single core. A
1775 repository of codes has been established at NERSC. A long-term goal is to make it easy for experimentalists
1776 to use these codes to compute Standard Model event rates for the processes they need.

1777 The **Distributed Computing and Facilities Infrastructures** subgroup looked at the growth trends in
1778 distributed resources as provided by the Open Science Grid, and the national high performance computing

1779 centers. Most of the computing by experiments is of the HTC type, but HPC centers could be used for
1780 specific work flows. Using existing computing centers could save smaller experiments from large investments
1781 in hardware and personnel. Distributed HTC has become important in a number of science areas outside
1782 HEP, but HEP is still the biggest user and must continue to drive the future computing development. HPC
1783 computing needs for theoretical physics will require an order of magnitude increase in capacity and capability
1784 at the HPC centers in the next five years, and two orders of magnitude in the next ten years.

1785 The **Networking** subgroup considered the implications of distributed computing on network needs, required
1786 R&D and engagement with the National Research and Education Networks (which carries most of our traffic).
1787 The group formulated a number of research questions that need to be answered before 2020. Expectations of
1788 network performance should be raised so that planning for network needs is on par with that for computing
1789 and storage. The gap between peak bandwidth and delivered bandwidth should be narrowed. Wide-area
1790 network performance should not be an insurmountable bottleneck in the next five to ten years as long as
1791 investments in higher performance links continue. However, there is uncertainty as to whether network costs
1792 will drop at the same rate as they have in the past.

1793 The **Software Development, Personnel and Training** subgroup proposed a number of recommendations
1794 to implement three main goals. The first goal is to use software development strategies and staffing models
1795 that result in software that is more widely useful to the HEP community. The second goal is to develop and
1796 support software that will run with optimal efficiency on future computer architectures. The third goal is to
1797 ensure that developers and users have the training necessary to deal with the increasingly complex software
1798 environments and computing systems that will be used in the future.

1799 The **Storage and Data Management** subgroup found that storage continues to be a cost driver for many
1800 experiments. It is necessary to manage the cost to optimize the science output from the experiment. Tape
1801 storage continues to be relatively inexpensive and should be utilized more within the storage hierarchy. Disk
1802 storage is likely to increase relatively slowly in capacity per unit cost, due to a shrinking consumer market
1803 and technology barriers. Operating distributed data management systems can be costly for experiments,
1804 and continued R&D in this area would benefit a number of experiments.

1805 To summarize, the challenging resource needs for the planned and proposed physics programs require efficient
1806 and flexible use of all resources. HEP needs both distributed HTC and HPC. Emerging experimental
1807 programs might consider a mix to fulfill demands. Programs to fund these resources need to continue.
1808 Sharing and opportunistic use help address resource needs, from all tiers of computing, eventually including
1809 commercial providers. There is increasing need for data-intensive computing in traditionally computation-
1810 intensive fields, including at HPC centers, for data analytics, combining simulations and observational data,
1811 etc.

1812 In order to satisfy our increasing computational demands, the field needs to make better use of advanced
1813 computing architectures. With the need for more parallelization, the complexity of software and systems
1814 continues to increase, impacting architectures for application frameworks, workload management systems,
1815 and also the physics code. We must develop and maintain expertise in all areas of the field, and re-engineer
1816 frameworks, libraries and physics codes. Unless corrective action is taken to enable us to take full advantage
1817 of the new hardware architectures, we could be frozen out of cost-effective computing solutions within 10
1818 years. There is a large code base that needs to be re-engineered, and we currently do not have enough people
1819 trained to do it.

1820 The continuing huge growth in observational and simulation data drives the need for continued R&D invest-
1821 ment in data management, data access methods, and networking. Continued evolution of data management
1822 and storage systems will be needed in order to take advantage of new network capabilities, ensure efficiency
1823 and robustness of the global data federations, and to contain the level of effort needed for operations.

1824 Significant challenges with data management and access remain, and research into these areas could continue
1825 to bring benefit across the frontiers. We expect solutions that will be based on content delivery approaches,
1826 dynamic data placement, and remote data access.

1827 Network reliability is essential for data-intensive distributed computing. Emerging network capabilities and
1828 data access technologies improve our ability to use resources independent of location. This will enable
1829 use of diverse compute resources. These include dedicated facilities, university computing centers, and
1830 opportunistic use of shared resources between PIs. They will expand to commercial clouds and eventually
1831 also make use of leadership-class HPC centers relevant for data-intensive computing. The computing models
1832 should treat networks as a resource that needs to be managed and planned for.

1833 Computing will be essential for progress in theory and experiment over the next two decades. The field
1834 continues to learn how to do more science with constrained resources, requiring us to be more flexible and
1835 perhaps tolerate higher levels of risk. The advances in computer hardware that we have seen in the past
1836 may not continue at the same rate in the future. The issues identified in this report require continuing
1837 attention. Addressing them will increase efficiency, reduce costs, and enable us to meet the experimental
1838 and theoretical goals identified through the Snowmass process.

1839 1.10 Communication, Education, and Outreach

1840 Broad societal support for particle physics research will be required to achieve the many scientific and
1841 technological goals identified by the U.S. particle physics community through the Snowmass process. Building
1842 and sustaining this support will require the particle physics community to unite behind a common plan that
1843 emerges from the Snowmass/P5 process and to communicate enthusiasm for the future of the field and its
1844 societal impacts to a wider audience of policy makers, opinion leaders, scientists in other fields, educators,
1845 and students.

1846 Federally supported research in particle physics and related fields has led to an impressive list of Nobel
1847 Prize-winning discoveries: the first detailed study of the cosmic microwave background, the discovery of
1848 neutrino masses and mixing (and earlier work on solar neutrinos many years earlier), the discovery of the
1849 accelerating expansion of the Universe, the understanding of the strong force, and the discovery that the
1850 CKM matrix explains CP violation. The most recent Nobel Prize was awarded for the Higgs boson, whose
1851 discovery in 2012 was made possible by scientific talent, technology and leadership from the United States.

1852 The American public is fascinated by these discoveries, and by the full breadth of current and future particle
1853 physics projects. The saga of the Large Hadron Collider and the Higgs boson discovery reached audience
1854 levels unprecedented for a particle physics event. Public lectures and other events on particle physics topics
1855 draw crowds. Milestones, discoveries and even proposals for projects in particle physics routinely make
1856 headlines.

1857 Translating this public excitement into greater support for the field requires existing communication, edu-
1858 cation and outreach (CE&O) activities to be augmented and enhanced. We need national coordination and
1859 training, additional resources, and the commitment by the particle physics community to convey consistent,
1860 coherent and compelling messages about the importance of particle physics research and its value to society.

1861 Many individuals, groups, and institutions in the U.S. particle physics community already reach out to
1862 members of the public, decision makers, teachers, and students through a wide variety of effective activities.
1863 However, there is room for improvement in the nationwide coordination of these activities, in the mobilization
1864 of the entire U.S. community to take part in the activities, and in efforts to use varied activities to convey
1865 consistent and compelling messages to stakeholders.

1866 The following is a survey of existing CE&O activities targeted at four audiences: policy makers and opinion
1867 leaders, scientists in other fields, the general public, and educators and students.

1868 **For policy makers and opinion leaders:** The U.S. particle physics community engages in a number of
1869 efforts to build support for research among policy makers and opinion leaders. User groups make annual visits
1870 to Washington, D.C. and, with scientific societies, conduct email and letter-writing campaigns at key points in
1871 the budget cycle. Scientists participate in Washington, D.C., events organized by the American Association
1872 for the Advancement of Science (AAAS), the American Physical Society (APS), the National User Facilities
1873 Organization (NUFO), and other organizations. Scientific and industrial societies, national laboratories, and
1874 individual scientists engage in direct advocacy with legislators. Scientists and media relations professionals
1875 at universities and labs work to place particle physics stories and physicists in influential media outlets.

1876 **For scientists in other fields:** Past and current outreach activities of the particle physics community
1877 targeted at colleagues in the broader science community include colloquia and seminars at university depart-
1878 ments and at national laboratories, and plenary sessions at APS and AAAS meetings. Particle physicists
1879 publish their results and author pedagogical review articles in journals such as *Science* and *Nature*. They
1880 write articles in *Scientific American*, *Popular Science* and similar magazines, write online articles and science
1881 blogs, publish popular science books, and author reports commissioned by labs and agencies. Examples of
1882 the last category include *Quantum Universe* and *Discovering the Quantum Universe*, prepared by HEPAP
1883 for the DOE and NSF.

1884 **For the general public:** Existing activities that reach the general public are extremely broad but have
1885 varying levels of support. The single most common activity is public talks. Scientists frequently participate
1886 in open houses and related events such as science festivals, lab and department tours, physics shows, alumni
1887 weekends, and workshops for the public. They contribute to external publications and shows by writing
1888 magazine articles and op-ed pieces in newspapers, participating as consultants to radio and television
1889 programs and movies, and working with the news media. They produce outreach materials such as books,
1890 brochures, posters, web-based materials, and multimedia products. They engage in social media activities
1891 such as blogs, Facebook pages, Twitter feeds, and the creation of YouTube videos.

1892 **For educators and students:** The particle physics community in the United States and abroad has
1893 succeeded in increasing student interest and achievement in STEM fields, including particle physics, through
1894 a variety of efforts. The 2013 APS Excellence in Physics Education Award was presented to leaders of Uni-
1895 versity of Illinois undergraduate physics education research. The QuarkNet long-term teacher development
1896 program has changed how many teachers view science and education by putting cosmic-ray detectors, online
1897 analysis tools and LHC data into their hands. Netzwerk Teilchenwelt adapted the QuarkNet model for
1898 students, teachers, and physicists in Germany. The International Particle Physics Outreach Group (IPPOG)
1899 sponsored 161 master classes in 37 countries in 2013, including 29 masterclasses in 9 countries in the Fermilab-
1900 based portion of the program. National laboratories run successful long-term education programs, many with
1901 particle physics content and partnerships with particle physics groups. The Contemporary Physics Education
1902 Project brings together physicists and educators to develop wall charts, posters, websites and activities.

1903 As a result of the Snowmass process, the community has recognized that more physicists must engage in
1904 CE&O activities. The quality and coordination of such activities must be improved in order to increase
1905 public support for the field, develop the next generation of physicists, and ensure scientifically literate and
1906 engaged citizens.

1907 A survey of 641 members of the particle physics community conducted in the spring of 2013 indicated that
1908 while about 60% of physicists engage in outreach to the general public and 50% reach K-12 teachers or
1909 students, only 30-35% engage in activities targeted to scientists in other fields or policy makers. The survey
1910 also identified the greatest barriers to participation in CE&O activities, including lack of time, little reward

1911 in career advancement and a lack of resources to communicate the broader societal impacts of particle physics
1912 research.

1913 At the Snowmass meeting, a number of prominent voices called for renewed commitment to CE&O:

- 1914 • “We must educate our representatives in Congress, our fellow citizens, the business community and
1915 the scientific agencies.”—D. Gross
- 1916 • “You are underselling yourselves. . . you are technology incubators for other fields of science.”—R. Roser
- 1917 • “The media missed the substantial impact of the U.S. on the Higgs discovery.”—J. Incandela
- 1918 • “You need to appeal to varied stakeholders to convince them that you do valuable science with a
1919 sensible plan. Illustrate the benefits of particle physics to society.”—G. Blazey

1920 The CE&O group developed the following goals, strategies, and recommendations with input from particle
1921 physicists and education and outreach professionals. The recommendations support a proactive, coordinated
1922 CE&O effort from the entire U.S. particle physics community.

1923 As overarching goals for U.S. particle physics communication, education and outreach, we recommend:

- 1924 1. Ensuring that the U.S. particle physics community has the resources necessary to conduct research
1925 and maintain a world leadership role.
- 1926 2. Ensuring that the U.S. public appreciates the value and excitement of particle physics.
- 1927 3. Ensuring that a talented and diverse group of students enters particle physics and other STEM careers,
1928 including science teaching.

1929 We recommend five-year CE&O implementation recommendations that cut across all audiences:

- 1930 1. Augment existing efforts with additional personnel and resources dedicated to nationwide coordination,
1931 training and support.
- 1932 2. Develop a comprehensive central communication, education and outreach resource for physicists, with
1933 initial content available before the end of the 2013/2014 P5 process.
- 1934 3. Provide communication training to the U.S. particle physics community.
- 1935 4. Work with DPF and HEPAP to develop a sustainable process for collecting statistics on workforce
1936 development and technology transfer and with APS to investigate a U.S. economic impact study for
1937 physics research that includes particle physics.

1938 We further recommend strategies for specific audiences:

1939 **For policy makers and opinion leaders:** (1) Empower and enable members of the particle physics
1940 community to communicate and advocate coherently, consistently, and effectively on behalf of their science.
1941 (2) Develop an enduring process to track, update, and disseminate statistics on the impact of particle
1942 physics on society. (3) Put informed third-party advocates to work raising the profile of and informing key
1943 stakeholders about the importance of particle physics, physics, and basic science to the United States.

1944 **For scientists in other fields:** (1) Foster more dialog and understanding between subfields of science. (2)
 1945 Identify areas of common cause and unite in support of them. (3) Develop consensus in our field that we
 1946 need to prioritize, buy into the mechanism of prioritization, and then support the resulting plan.

1947 **For the general public:** (1) Engage the public in a wide range of outreach activities. (2) Make the public
 1948 aware of direct and indirect applications of research, both historical and potential. (3) Communicate the
 1949 role and stories of U.S. physicists in particle physics, particularly in major discoveries and in the context of
 1950 our international collaborations.

1951 **For educators and students:** (1) Directly engage with students and educators. Invite educators and
 1952 students into our unique community. (2) Offer long-term professional development and training opportunities
 1953 for educators (including pre-service educators), aligned with current and appropriate standards and enabling
 1954 educators to explore best-practice teaching methods. Make an effort to collaborate with local schools of
 1955 education whenever possible. (3) Create learning opportunities for students of all ages, including classroom,
 1956 out-of-school and online activities that allow students to explore particle physics to construct their own
 1957 understanding and develop the skills and habits of mind necessary to perform research.

1958 **Resources:** An overarching recommendation that supports all goals and strategies is the augmentation
 1959 of existing efforts with additional personnel and resources dedicated to nationwide coordination, training,
 1960 and support for particle physics education, outreach, and communication activities. Such a team would
 1961 enhance existing efforts and spearhead new initiatives, such as the development of a comprehensive central
 1962 communication, education and outreach resource for physicists, the development of sustainable methods
 1963 to collect statistics on workforce development and technology transfer, materials designed to inform the
 1964 public about direct and indirect applications of particle physics, the creation of professional development
 1965 opportunities for educators, and new learning opportunities for students of all ages.

1966 1.11 Conclusion

1967 Here we recapitulate the Conclusion given previously in our Executive Summary. With the completion of
 1968 the Standard Model, particle physicists now turn their attention to still deeper questions about the nature of
 1969 matter and the constituents of the universe. This report proposes an ambitious array of new experiments. We
 1970 consider it realistic to carry out these experiments through a long-term plan and through global partnerships.
 1971 Particle physicists have been the pioneers of large-scale scientific projects. We have constructed facilities of
 1972 unprecedented scale, including the Tevatron and the Large Hadron Collider, through decades-long programs
 1973 requiring world-wide collaboration. These led to discoveries that are the foundation of our current success.

1974 Several strategic goals have emerged from the Snowmass study.

- 1975 • Probe the highest possible energies and distance scales with the existing and upgraded Large Hadron
 1976 Collider and reach for even higher precision with a lepton collider; study the properties of the Higgs
 1977 boson in full detail.
- 1978 • Develop technologies for the long-term future to build multi-TeV lepton colliders and 100 TeV hadron
 1979 colliders.
- 1980 • Execute a program with the U.S. as host that provides precision tests of the neutrino sector with an
 1981 underground detector; search for new physics in quark and lepton decays in conjunction with precision
 1982 measurements of electric dipole and anomalous magnetic moments.

- 1983 • Identify the particles that make up dark matter through complementary experiments deep under-
1984 ground, on the Earth's surface, and in space, and determine the properties of the dark sector.
- 1985 • Map the evolution of the universe to reveal the origin of cosmic inflation, unravel the mystery of dark
1986 energy, and determine the ultimate fate of the cosmos.
- 1987 • Invest in the development of new, enabling instrumentation and accelerator technology.
- 1988 • Invest in advanced computing technology and programming expertise essential to both experiment and
1989 theory.
- 1990 • Carry on theoretical work in support of experimental projects and to explore new unifying frameworks.
- 1991 • Invest in the training of physicists to develop the most creative minds to generate new ideas in theory
1992 and experiment that advance science and benefit the broader society.
- 1993 • Increase our efforts to convey the excitement of our field to others.

1994 In pursuit of these projects, we have developed a community that links together scientists from all regions
1995 of the world pursuing common goals. Our community is ready and eager to carry out the next steps in
1996 humankind's quest to understand the basic workings of the universe.